

Magnetic fields in the ISM & their effect on star formation

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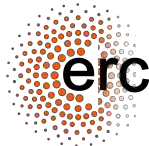
with Matthew Bate & Daniel Price

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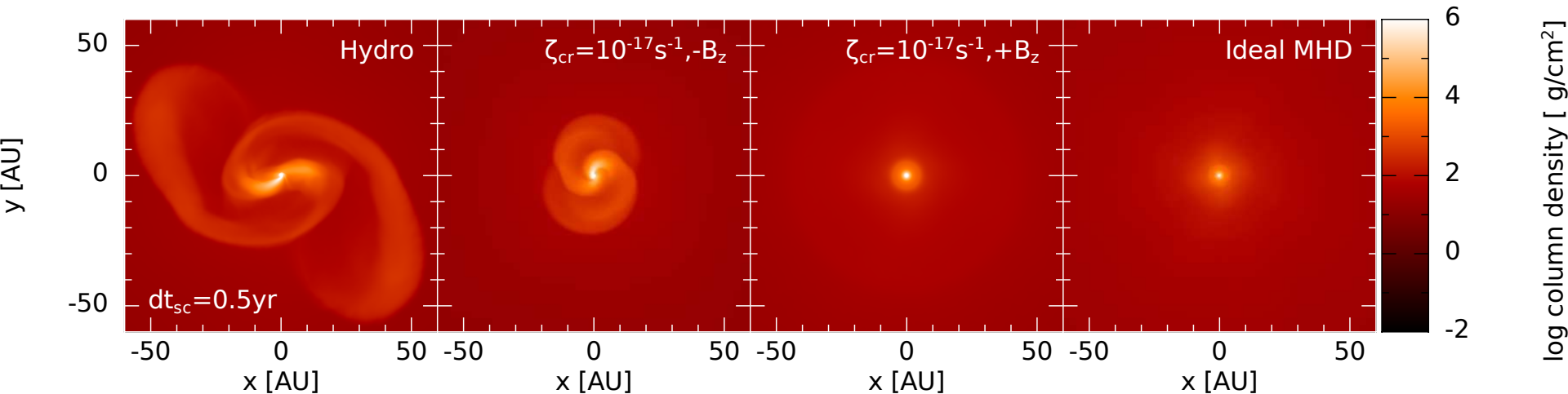
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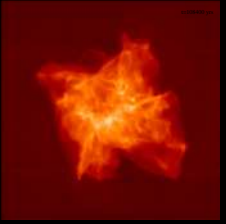
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Disc formation: effect of magnetic fields

➤ Using idealised initial conditions:

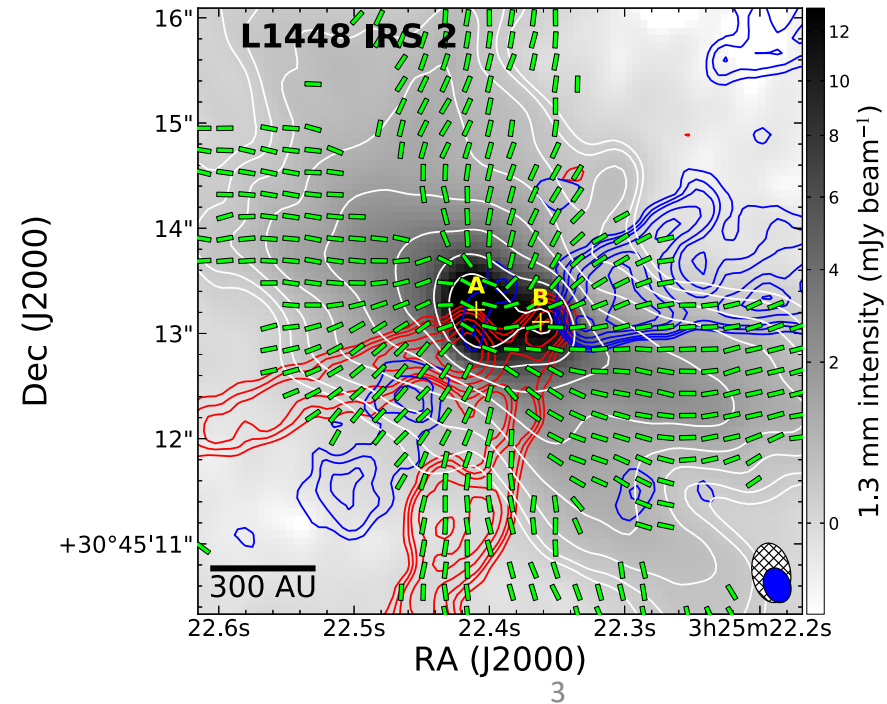
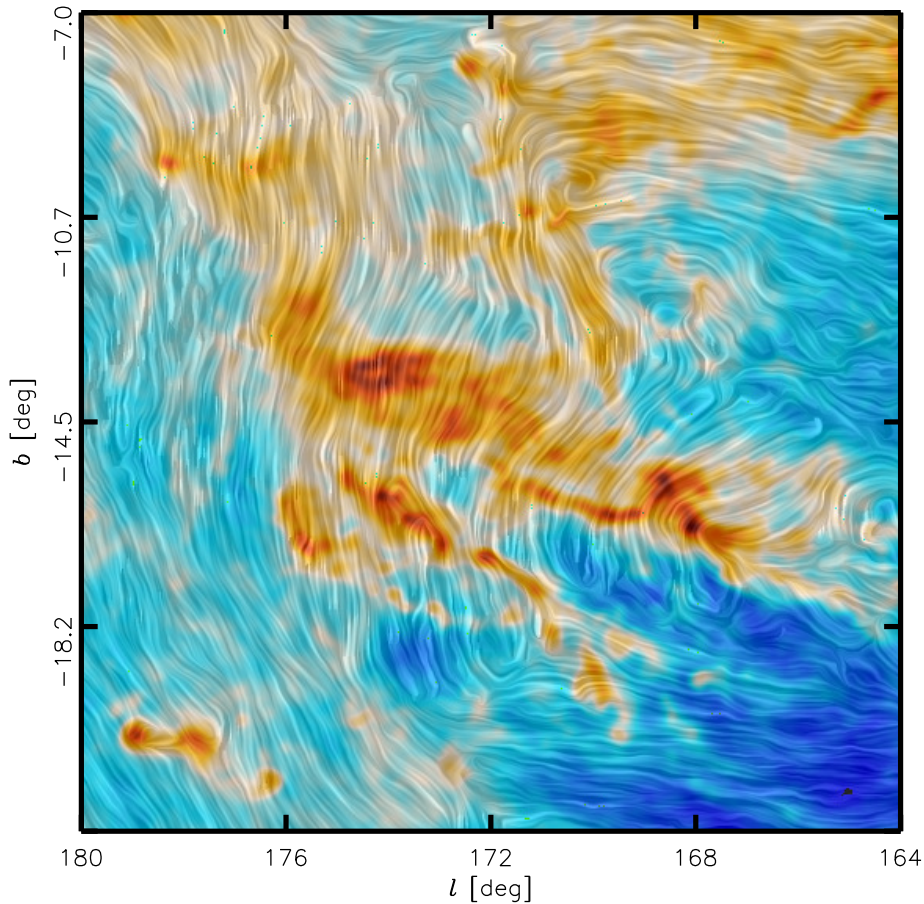


- Hydro:
forms a 50au disc early during the first hydrostatic core phase
- Non-ideal (Ohmic resistivity, ambipolar diffusion & the Hall effect)
 - $-B_z$: forms a 25au disc during the first hydrostatic core phase
 - $+B_z$: forms a 1au disc by the stellar core phase
- Ideal:
never forms a rotationally supported disc
(the Magnetic Braking Catastrophe; Allen, Li & Shu, 2003)



Disc formation:

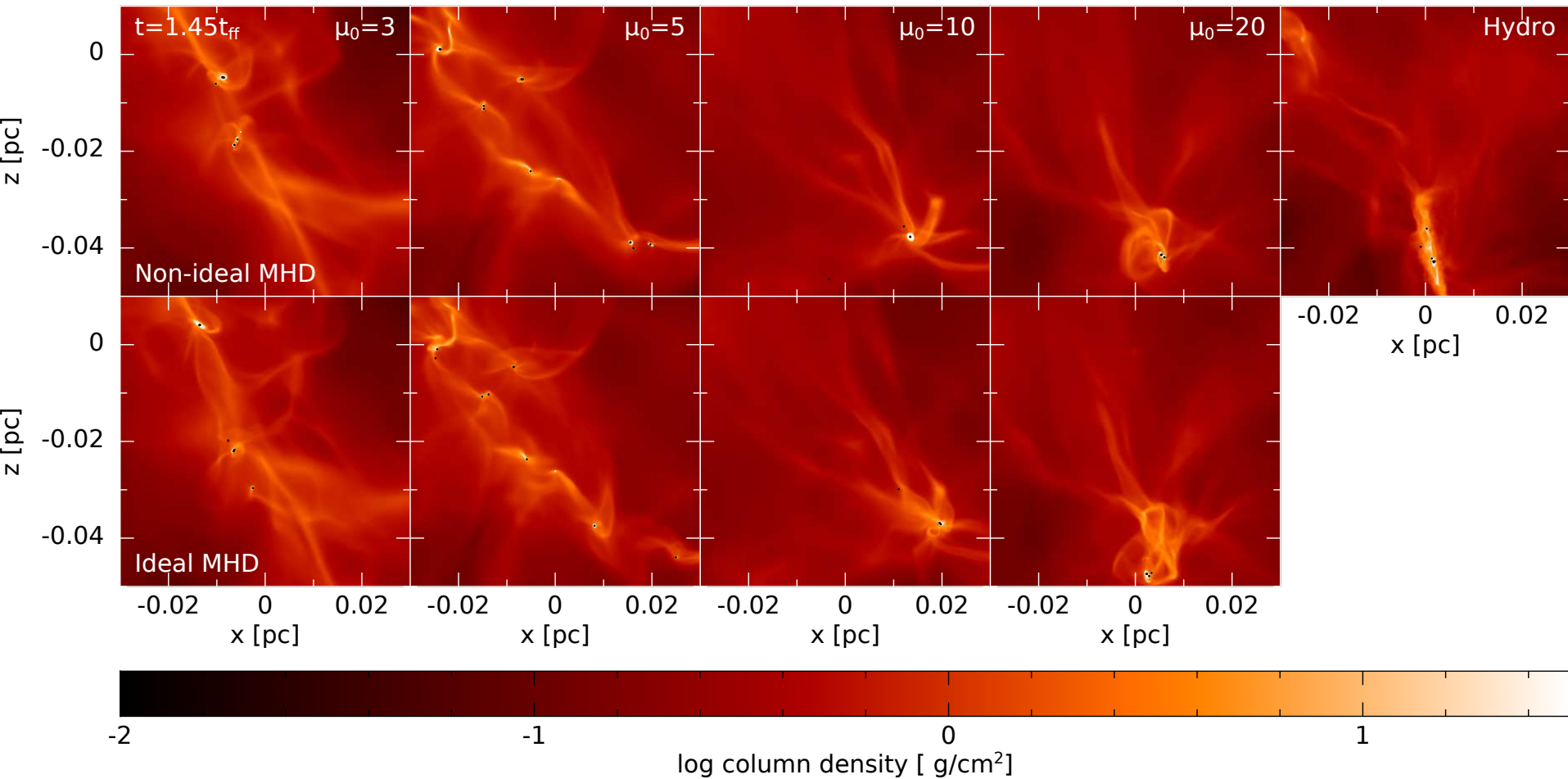
- Large-scale magnetic fields are perpendicular to dense structures
- Large-scale magnetic fields are parallel to low-density structures
- Small-scale magnetic fields get pinched into dense regions, creating an hour-glass shape



Cluster formation: effect of magnetic fields

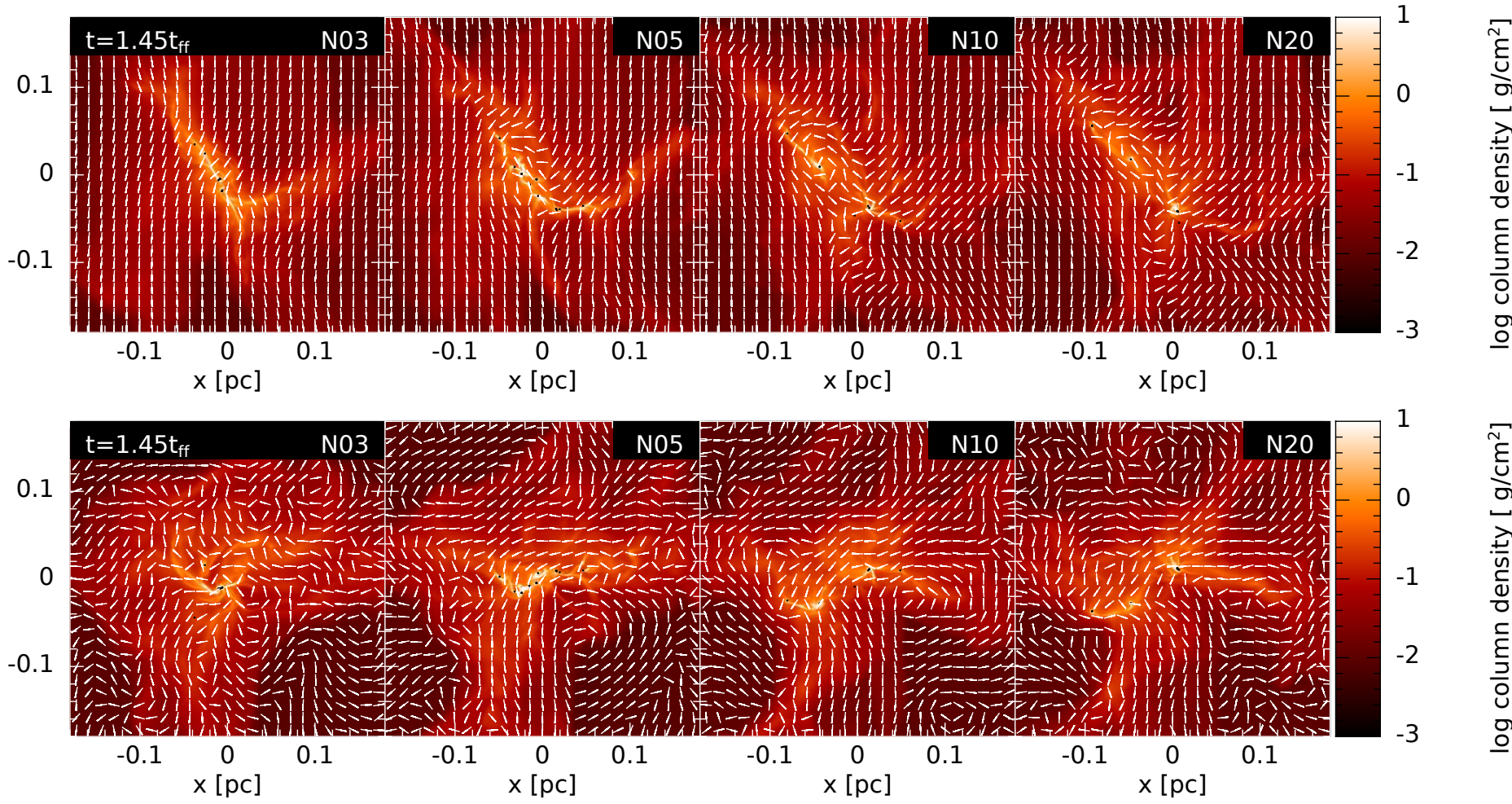


Decreasing magnetic field strength



Cluster Formation: Magnetic field lines

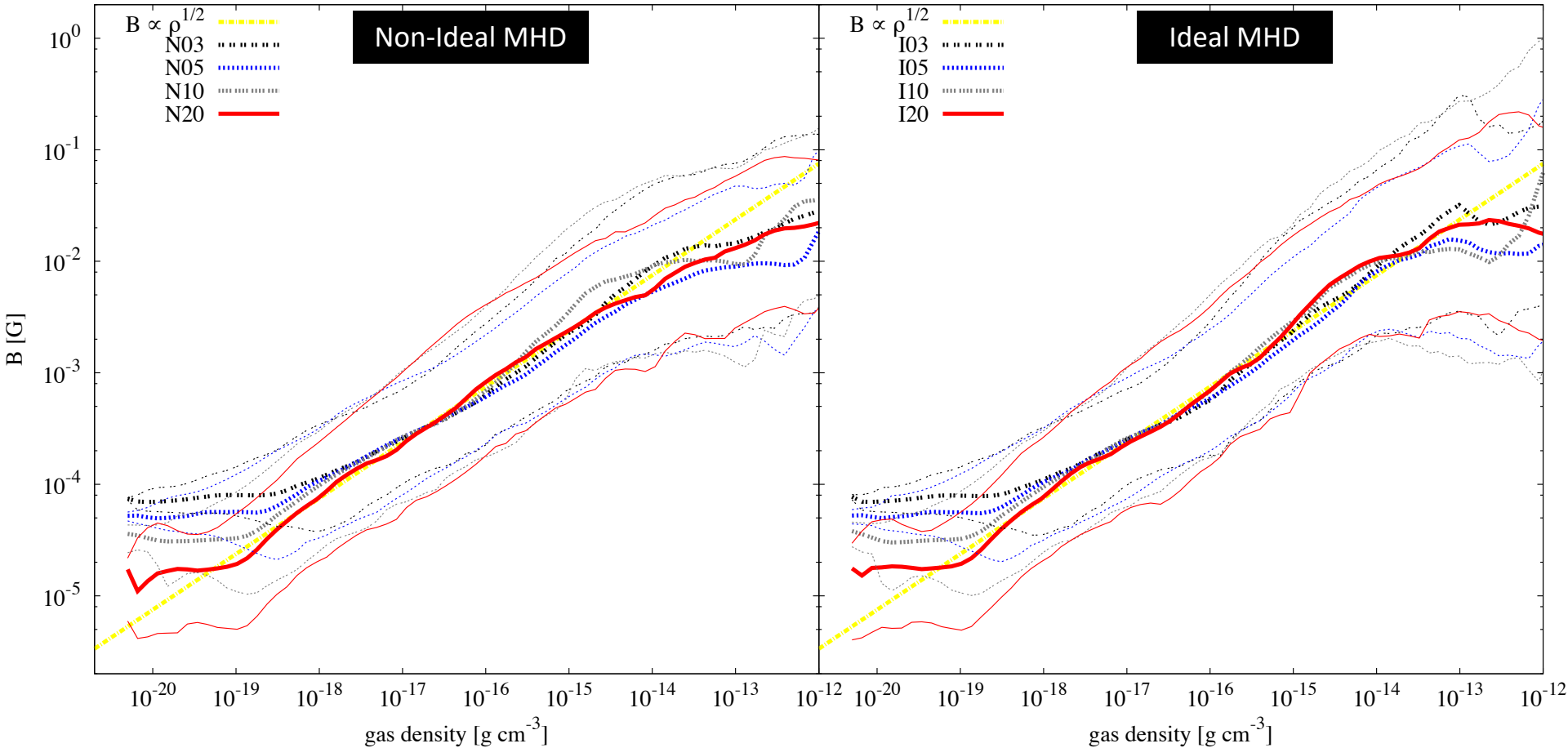
Decreasing magnetic field strength





Cluster Formation: Star forming regions

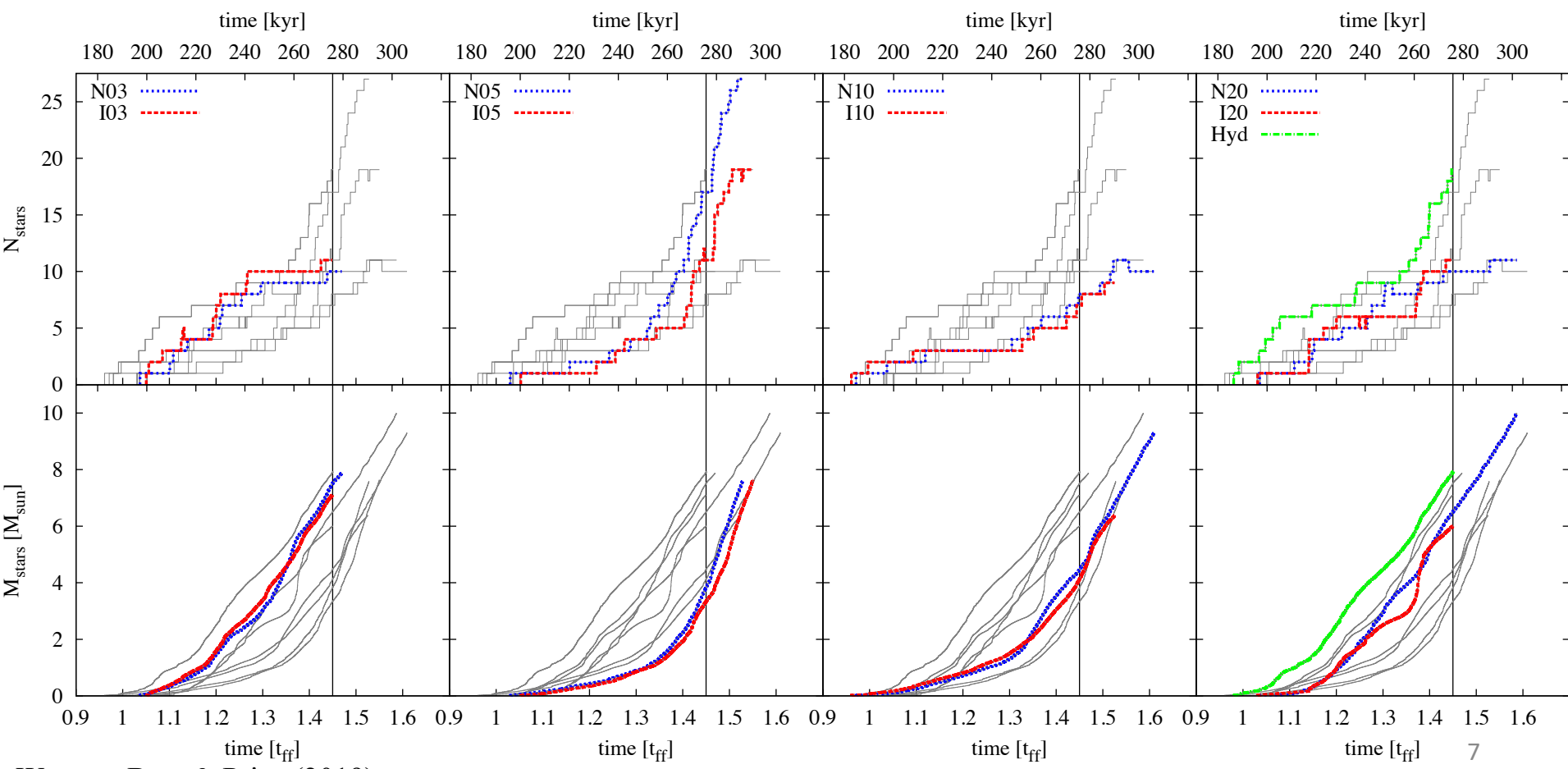
- Star forming regions have a wide range of initial magnetic field strengths, that are approximately independent of the global environment



Cluster Formation: Stellar Mass

- No trend when stars form
- Excluding N03 & I03, there is more mass in stars with weaker initial magnetic field strengths

Decreasing magnetic field strength \longrightarrow





Conclusions

- Isolated stars:
 - Magnetic fields play an important role in determining formation of the star and its discs
 - Non-ideal MHD is instrumental in disc formation
- Star cluster formation:
 - Ideal MHD affects large-scale structures, while non-ideal MHD affects small-scale structures
 - No trend between magnetic fields strength and number of stars
 - Loose trend between magnetic field strength and total mass of stars

