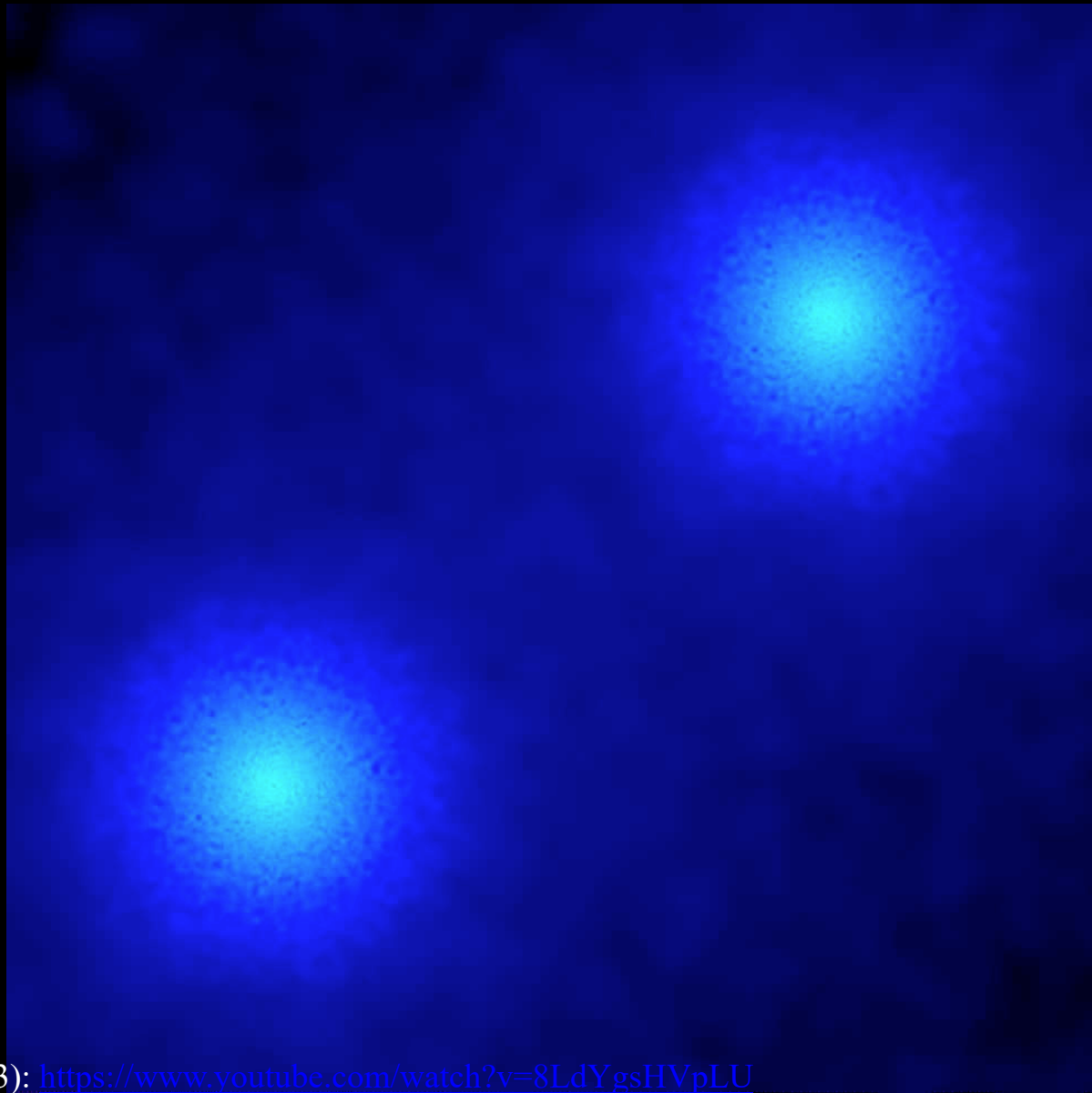




Numerical Hydrodynamics II



Example: Astrophysics: Galaxy merger





Complex Calculation: Required Components

Re-cap

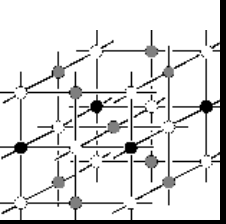
➤ To solve any system numerically, we require

A method to divide the region (e.g. grids)

A method to describe the evolution of the region
(i.e. the set of fluid dynamics equations)

A method to describe the edge of the region (i.e. boundary conditions)

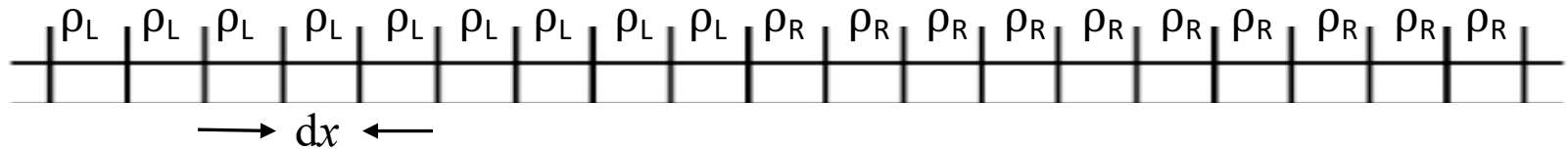
The initial properties of the system (i.e. initial conditions)



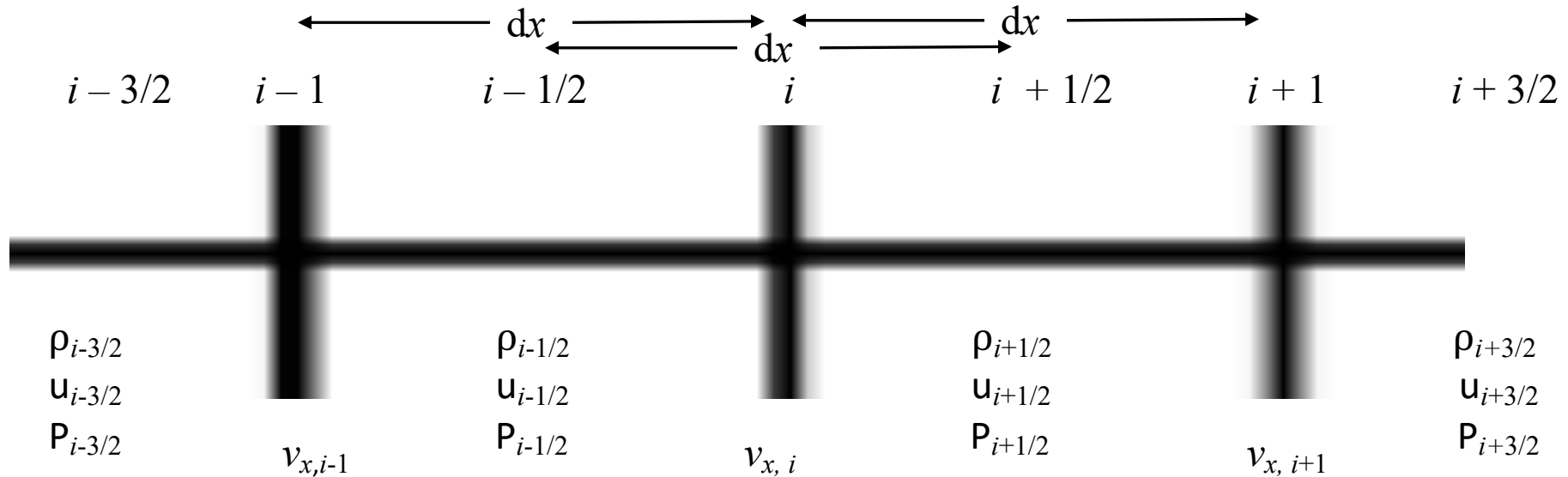
Defining your problem: Defining quantities

Re-cap

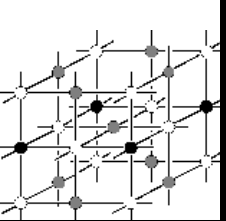
- Eulerian grid: grid of constant spacing



- A few cells:



- Scalars are calculated at *cell-centre*
- Vectors are calculated at *cell-interface*



Fluid equations: Continuum vs 1D-Numerical

Re-cap

➤ Continuum equations

Continuity equation: $\frac{D\rho}{Dt} = -\rho \nabla \cdot \mathbf{v}$

Equation of motion: $\frac{D\mathbf{v}}{Dt} = -\frac{1}{\rho} \nabla P$

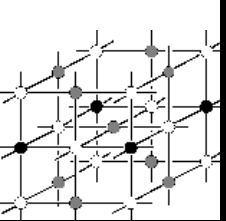
Energy equation: $\frac{Du}{Dt} = -\frac{P}{\rho} \nabla \cdot \mathbf{v}$

Equation of state: $P = (\gamma - 1) \rho u$

➤ There exists various schemes (e.g. donner-cell) that stabilise the code

$$i - 3/2 \quad i - 1 \quad i - 1/2 \quad i \quad i + 1/2 \quad i + 1 \quad i + 3/2$$





Fluid equations: Continuum vs 1D-Numerical

Re-cap

➤ Discrete equations in Eulerian form:

$$v_{x,i}^{n+\frac{1}{2}} = v_{x,i}^{n-\frac{1}{2}} - dt \left(\frac{2}{\rho_{i+\frac{1}{2}}^n + \rho_{i-\frac{1}{2}}^n} \frac{P_{i+\frac{1}{2}}^n - P_{i-\frac{1}{2}}^n}{dx} + v_{x,i}^{n-\frac{1}{2}} f(v) \right)$$

$$\rho_{i+\frac{1}{2}}^{n+1} = \rho_{i+\frac{1}{2}}^n - dt \left(\rho_{i+\frac{1}{2}}^n \frac{v_{x,i+1}^{n+\frac{1}{2}} - v_{x,i}^{n+\frac{1}{2}}}{dx} + \frac{v_{x,i+1}^{n+\frac{1}{2}} + v_{x,i}^{n+\frac{1}{2}}}{2} f(\rho) \right)$$

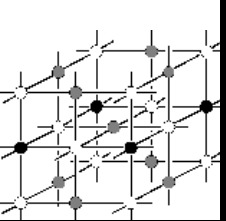
$$u_{i+\frac{1}{2}}^{n+1} = u_{i+\frac{1}{2}}^n - dt \left(\frac{P_{i+\frac{1}{2}}^n}{\rho_{i+\frac{1}{2}}^n} \frac{v_{x,i+1}^{n+\frac{1}{2}} - v_{x,i}^{n+\frac{1}{2}}}{dx} + \frac{v_{x,i+1}^{n+\frac{1}{2}} + v_{x,i}^{n+\frac{1}{2}}}{2} f(u) \right)$$

$$P_{i+\frac{1}{2}}^{n+1} = (\gamma - 1) \rho_{i+\frac{1}{2}}^{n+1} u_{i+\frac{1}{2}}^{n+1}$$

➤ $f(a)$ represents the Lagrangian part of the derivative, and can be first, second, third, ..., order

$i - 3/2 \quad i - 1 \quad i - 1/2 \quad i \quad i + 1/2 \quad i + 1 \quad i + 3/2$

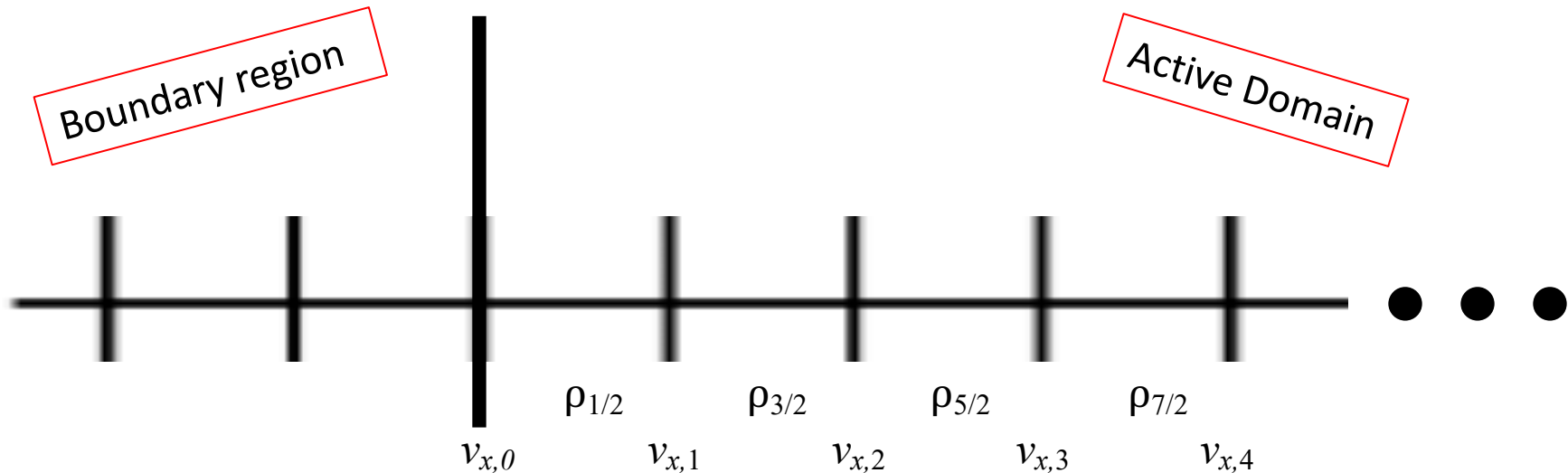


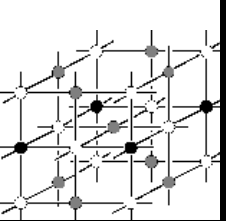


Boundaries

Re-cap

- We almost have enough information to run a simple simulation, but what happens at the edge of the simulation?
- Similar to solving differential equations, boundary conditions are required (e.g.)
 - Fixed / Inflow
 - Outflow
 - Reflective
 - Periodic

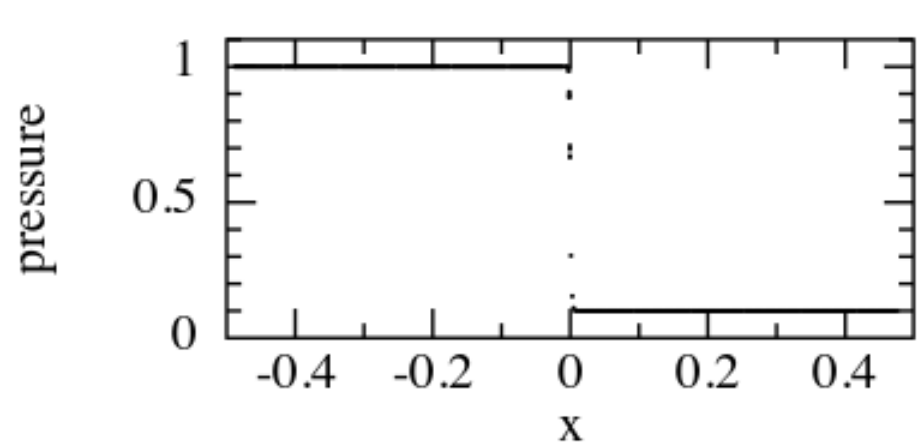
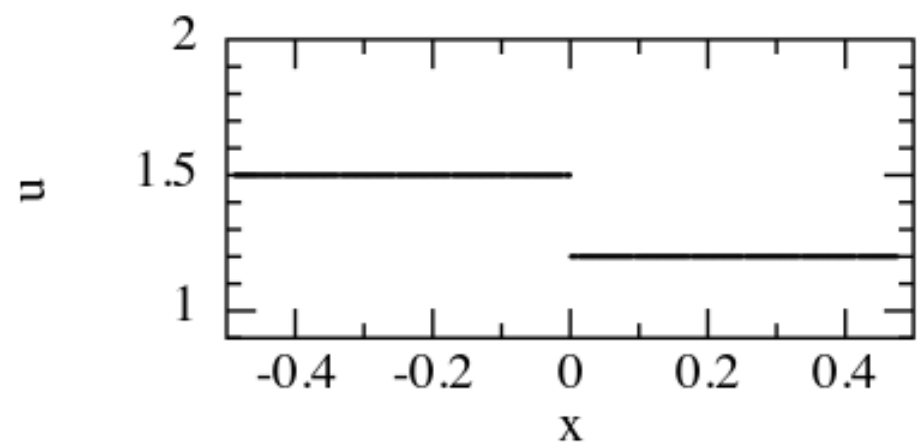
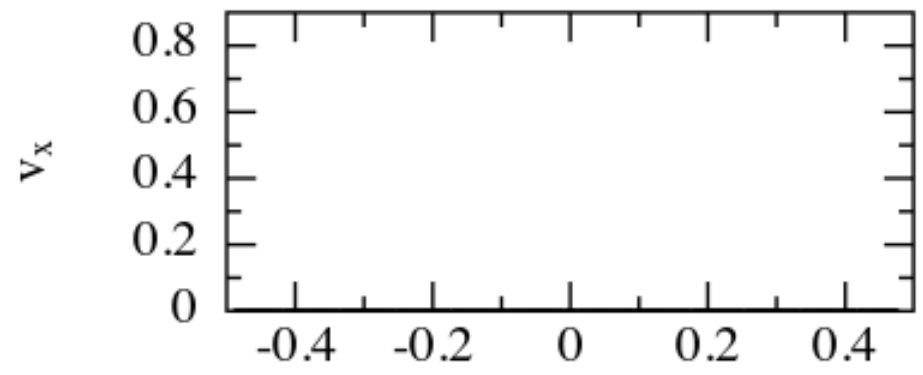
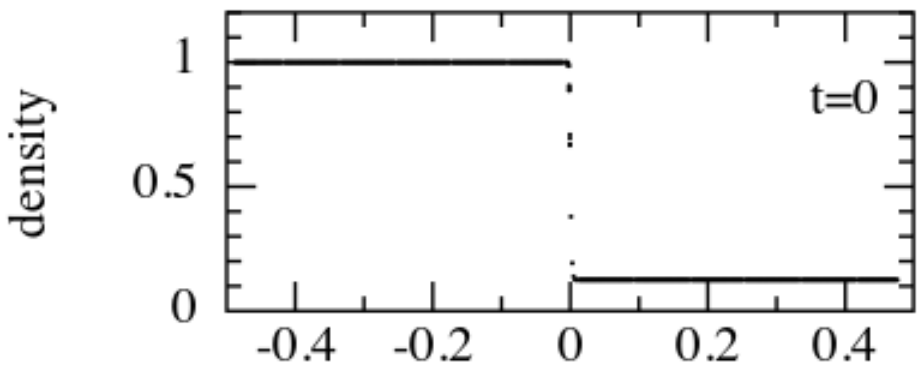


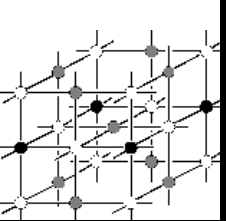


Initial conditions: Sod Shock

Re-cap

- Initial conditions for the Sod Shock
- Boundary Conditions: fixed (since we stop the problem before the shock hits the walls)

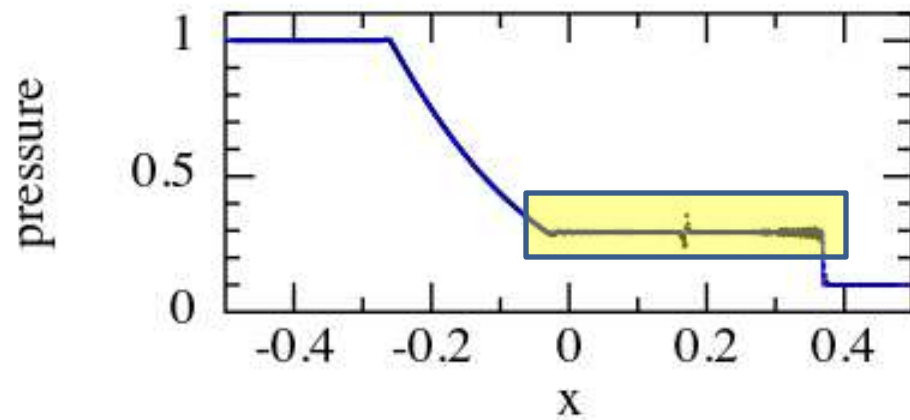
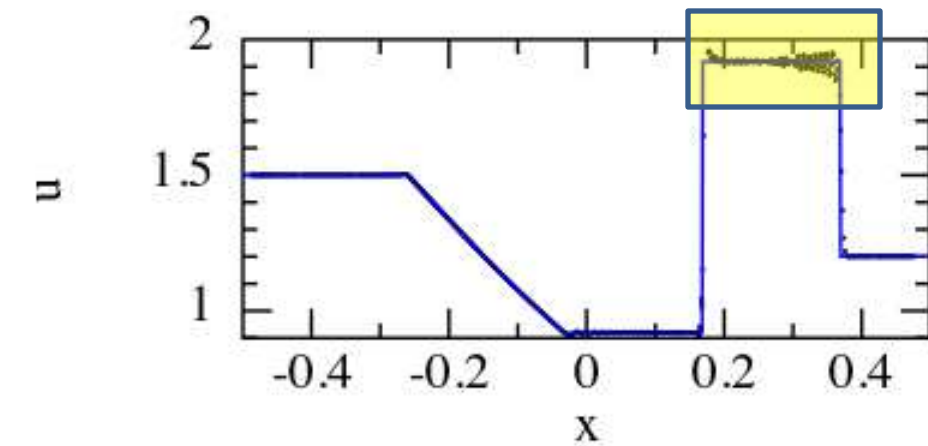
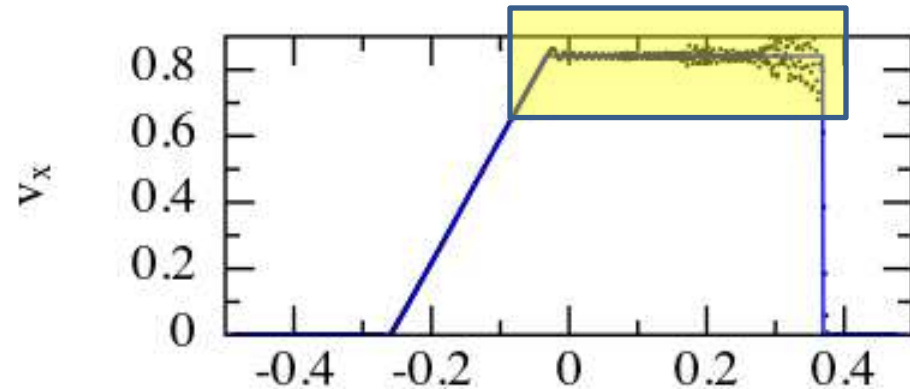
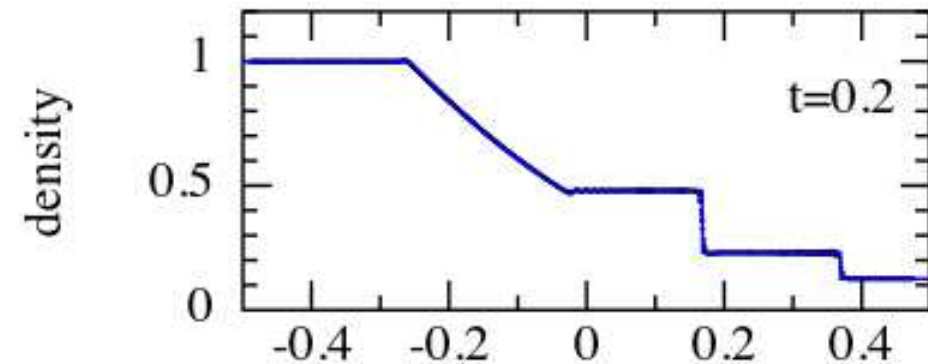


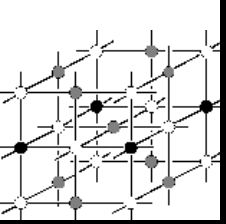


Sod Shock Evolution

Re-cap

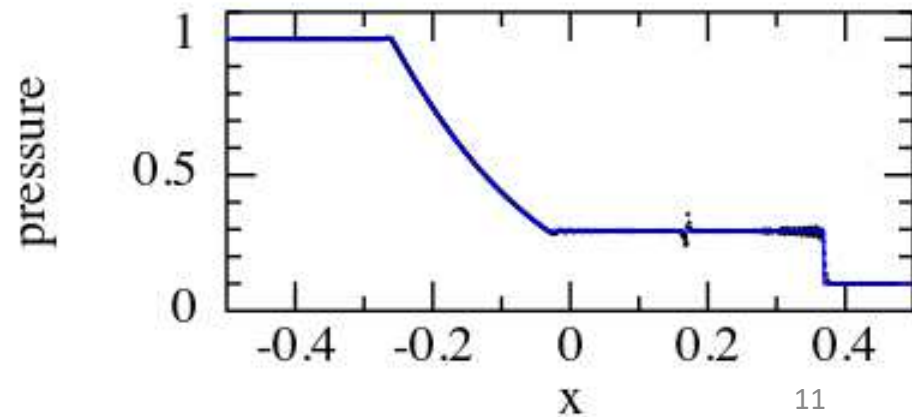
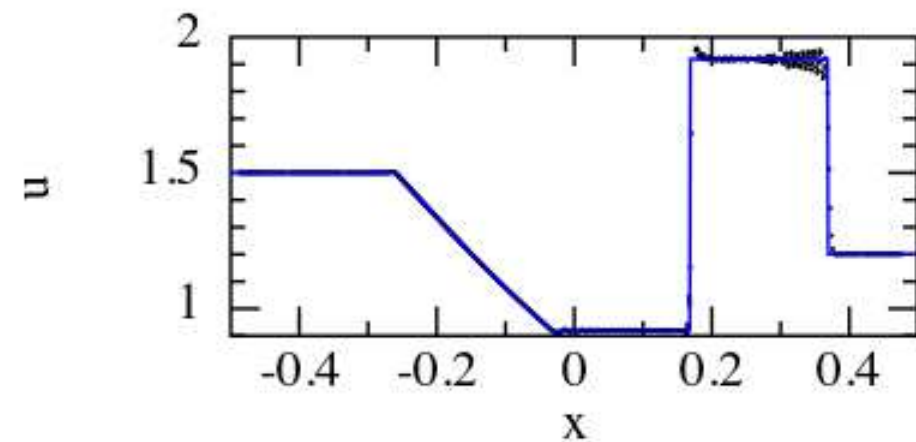
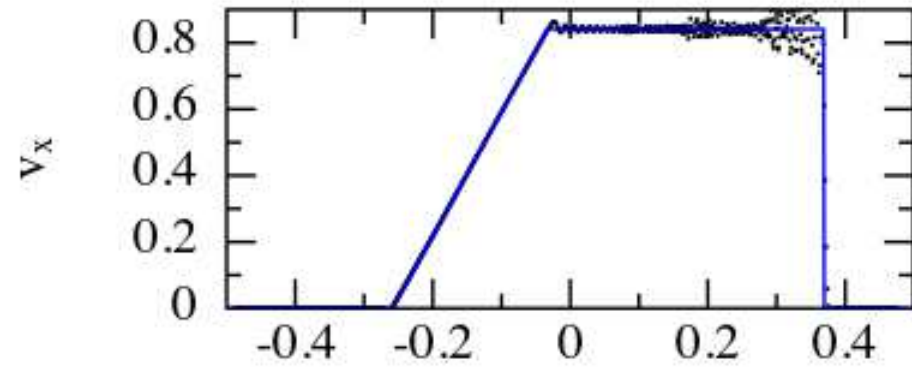
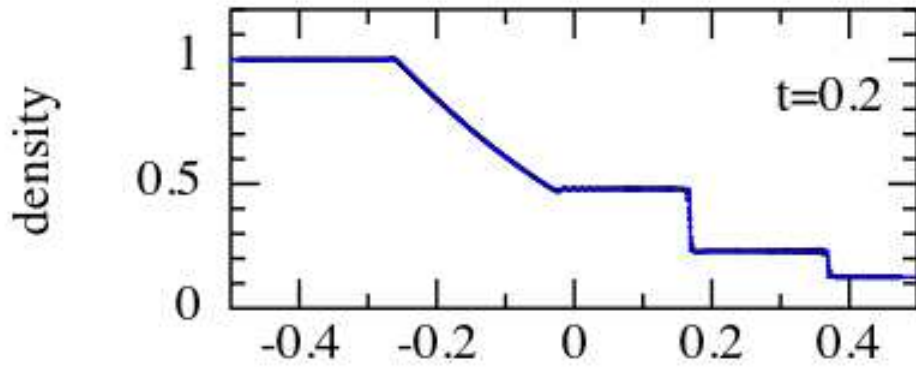
- Ringing and instabilities occur at the shock wave dampen as the shock propagates

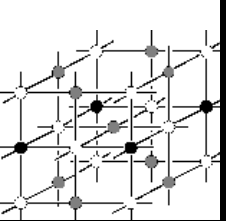




Sod Shock: *Structure of shock*

- Physical acceptable solution contains three distinct waves:
 - shock wave
 - contact discontinuity
 - rarefaction wave

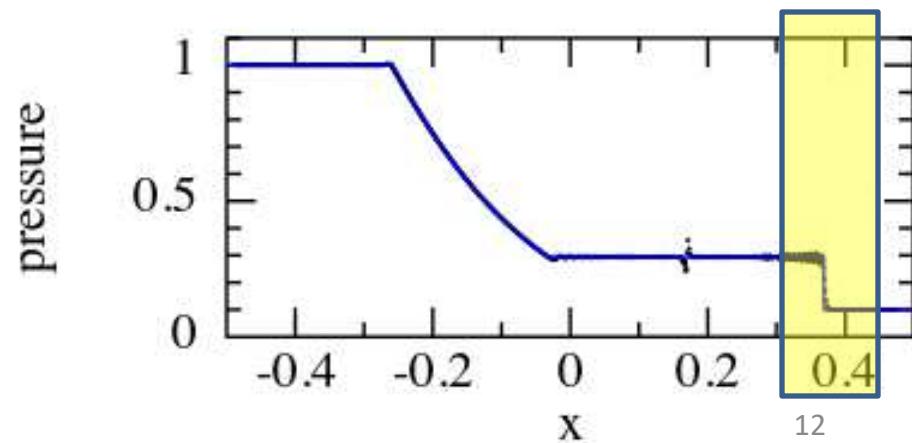
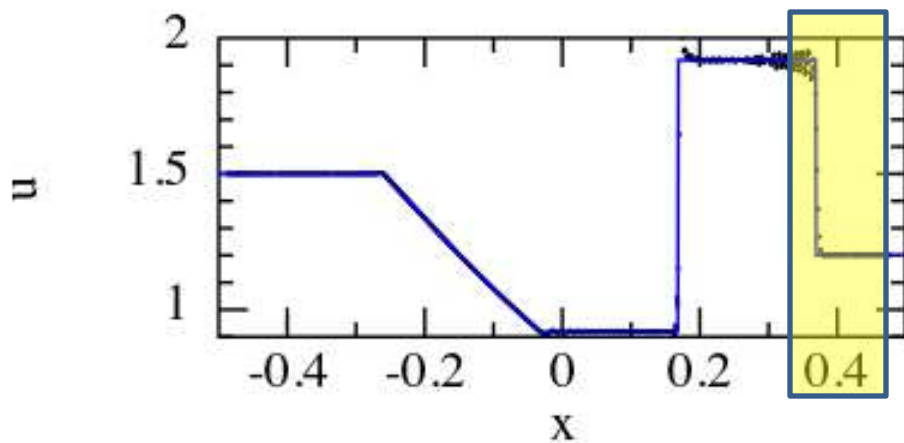
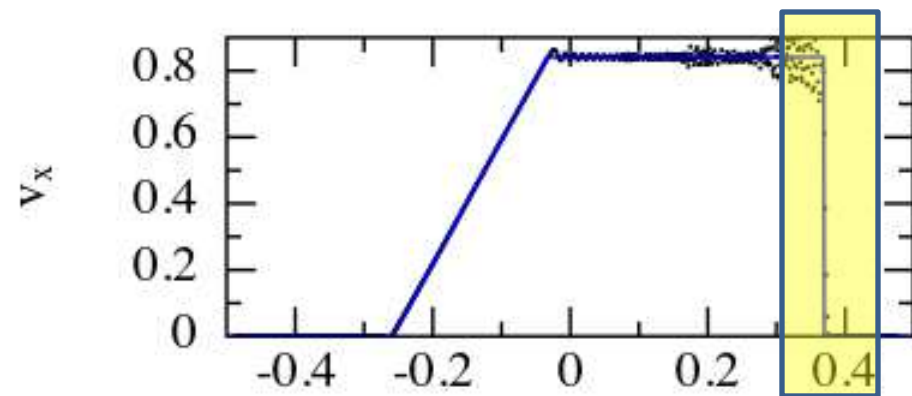
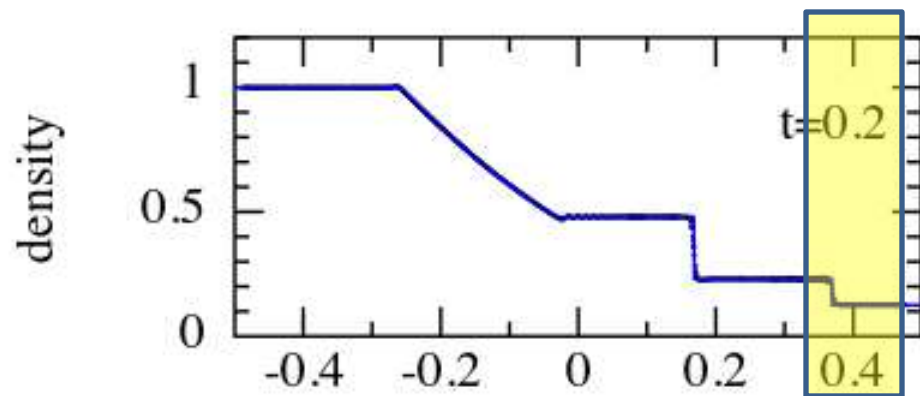


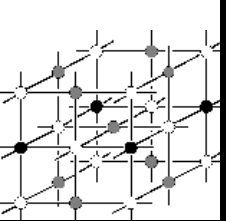


Sod Shock: *Structure of shock*

➤ Shock Wave

- Strong discontinuity in density, pressure and fluid velocity
- Supersonic movement caused by a strong pressure or velocity gradient

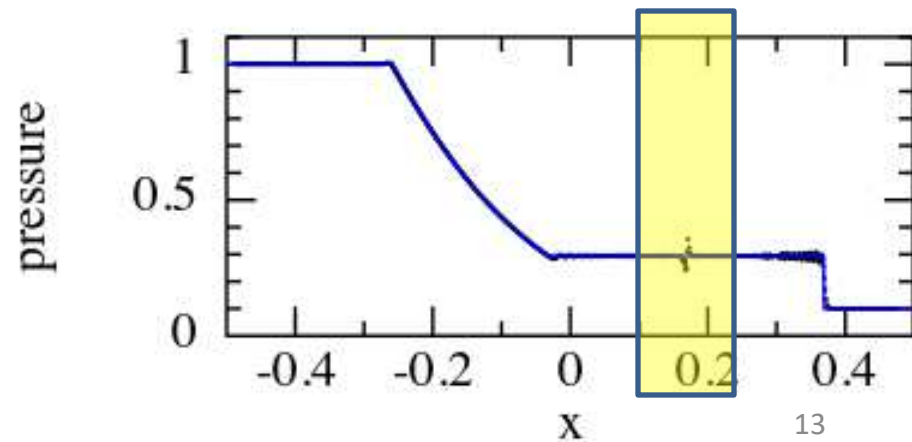
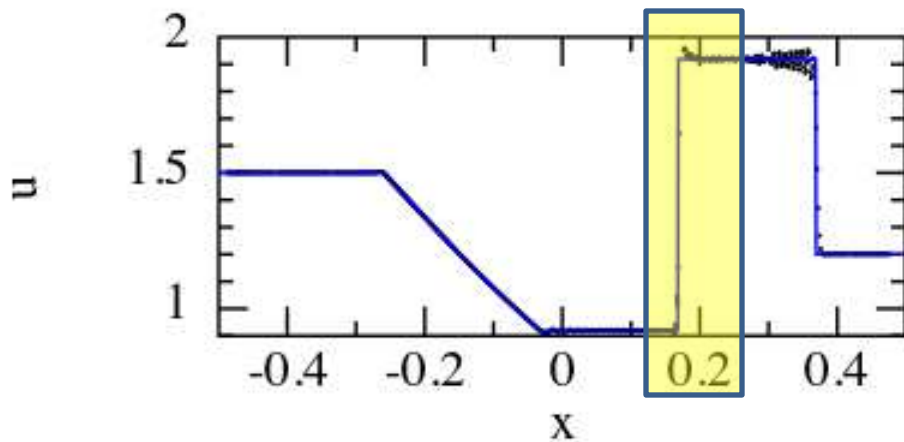
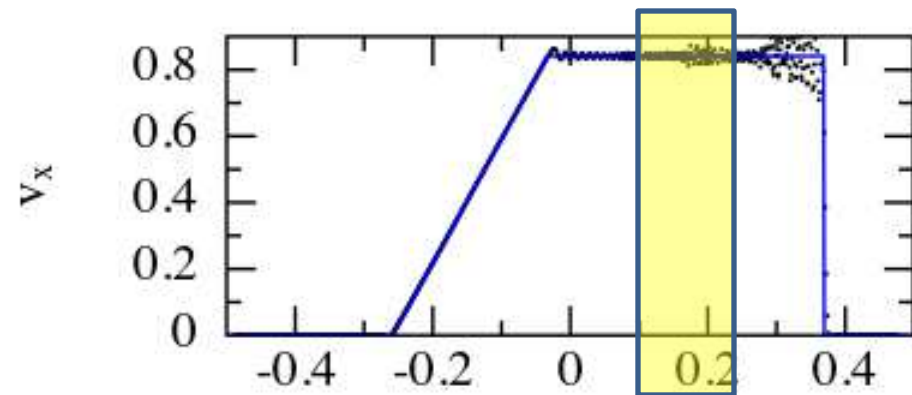
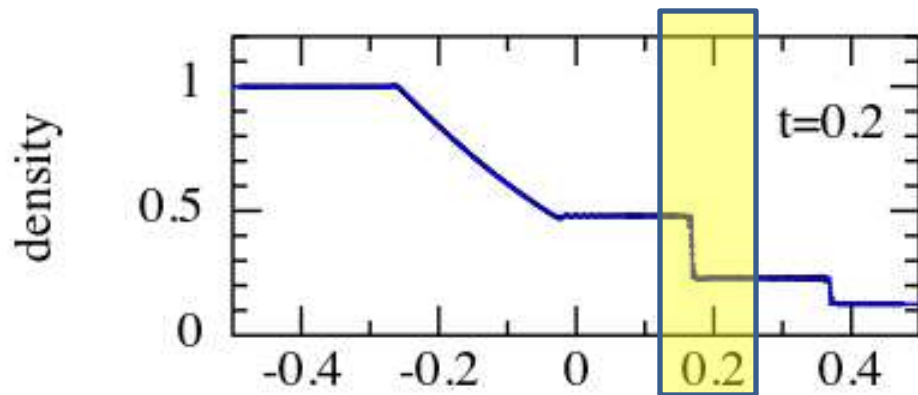


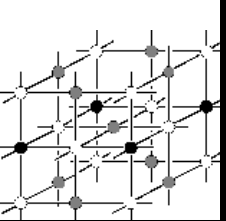


Sod Shock:

Structure of shock

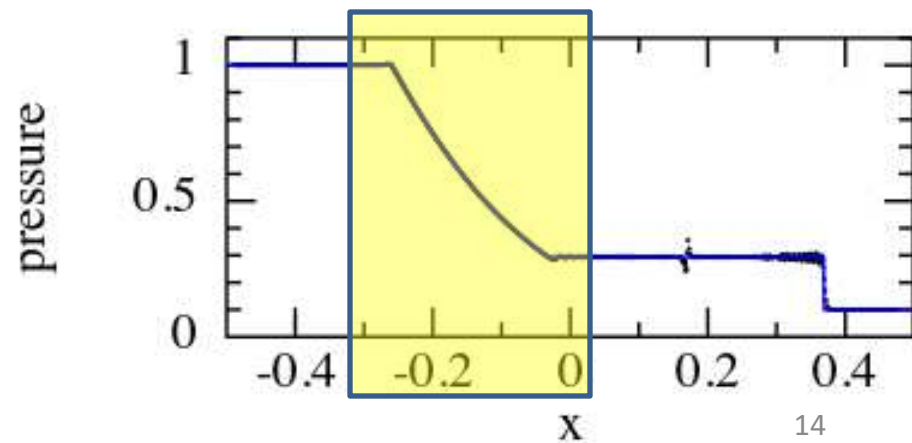
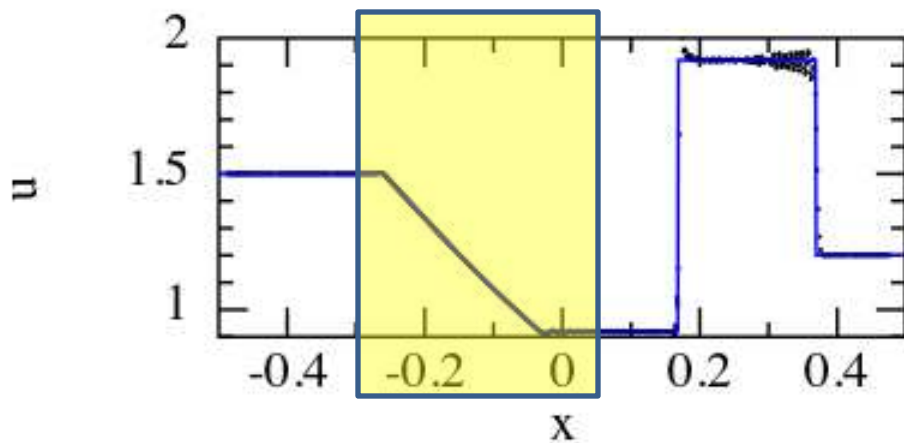
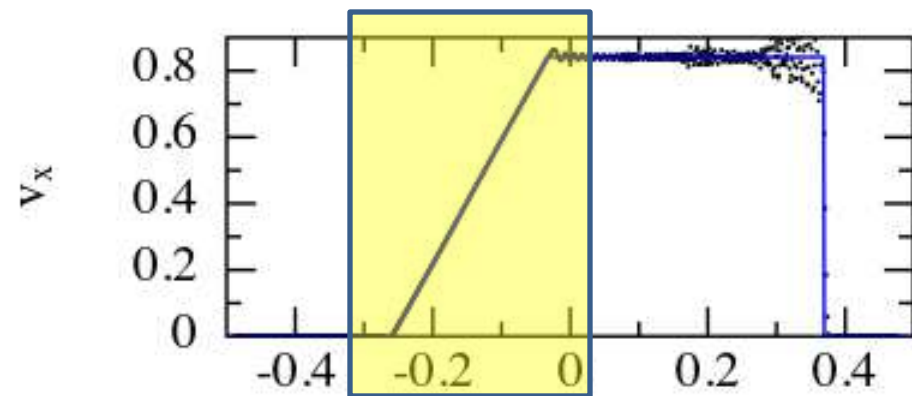
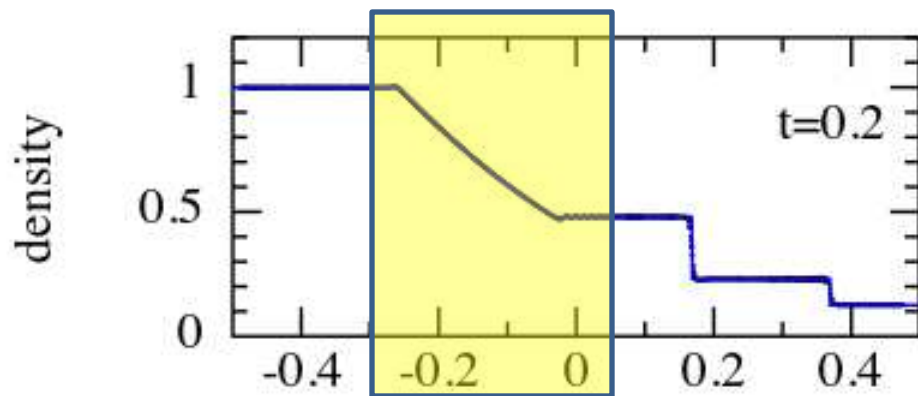
- Contact discontinuity
 - Discontinuity in density; pressure and fluid velocity are constant
 - Moving with the local fluid velocity

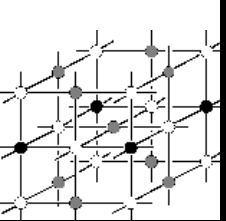




Sod Shock: *Structure of shock*

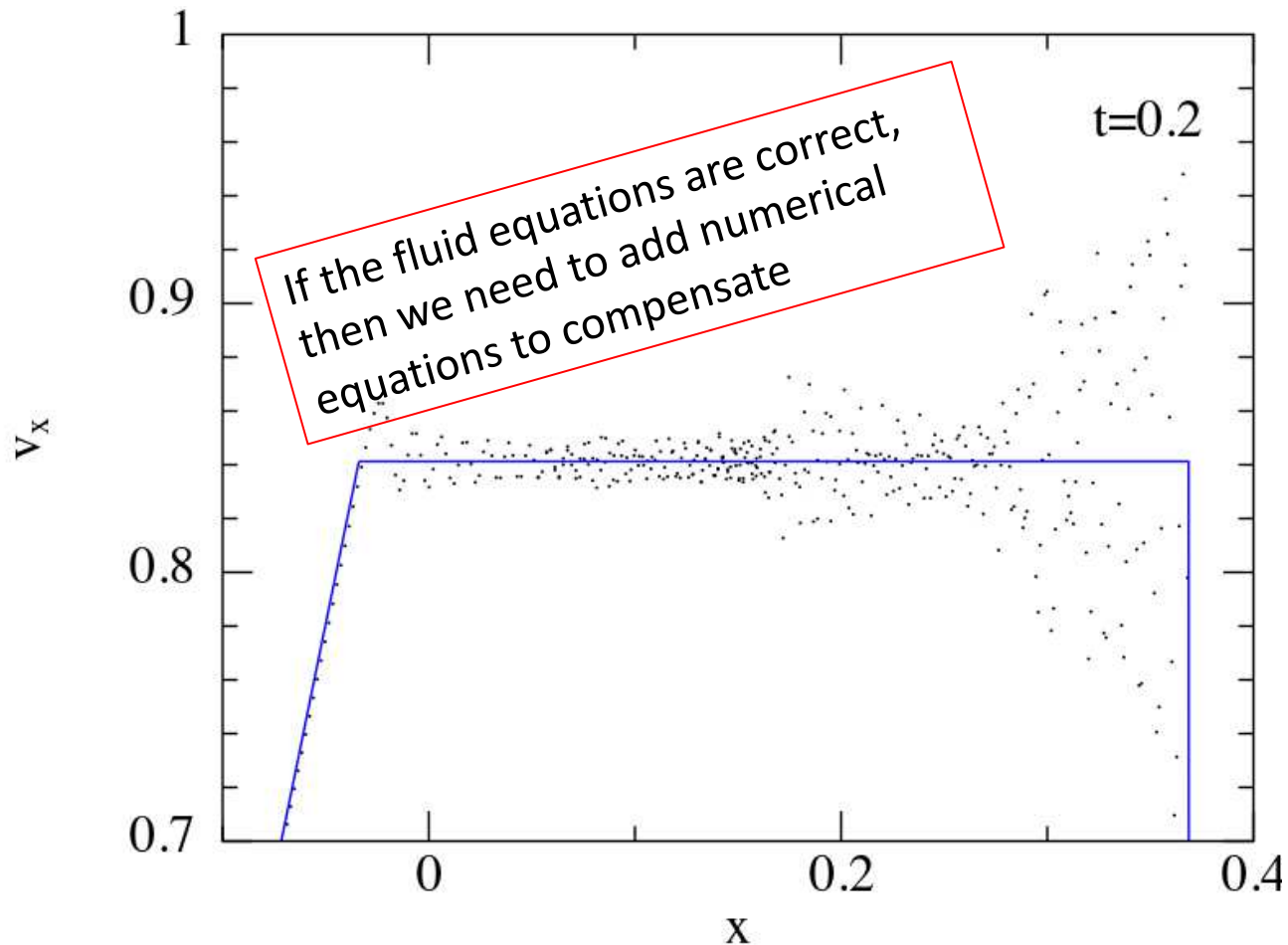
- Rarefaction wave
 - Continuous change in density, pressure and fluid velocity
 - Moving with the sound speed relative with respect to the local fluid velocity

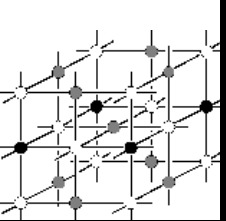




Sod Shock Evolution

- Ringing and instabilities occur at the **shock wave** & dampen as the shock propagates
 - Numerical methods often have difficulty resolving sharp discontinuities
 - The algorithms typically overpredict in one cell, then underpredict in the next





Sod Shock:

Artificial terms: Artificial viscosity

- Numerical algorithms are required for stability
- The form and parameterisation of these requires careful consideration to suppress numerical instabilities but not physical instabilities
- Modify the velocity by adding in an artificial pressure term:

$$v_{x,i}^{n+\frac{1}{2}} = v_{x,i}^{n-\frac{1}{2}} - dt \left[\frac{2}{\rho_{i+\frac{1}{2}}^n + \rho_{i-\frac{1}{2}}^n} \frac{\left(P_{i+\frac{1}{2}}^n + q_{i+\frac{1}{2}}^n \right) - \left(P_{i-\frac{1}{2}}^n + q_{i-\frac{1}{2}}^n \right)}{dx} + v_{x,i}^{n-\frac{1}{2}} f(v) \right]$$

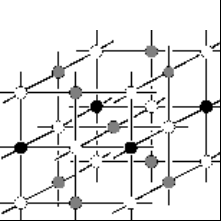
where

$$q_{i+\frac{1}{2}}^n = \begin{cases} \alpha^2 \rho_{i+\frac{1}{2}}^n \Delta v_i^2 & \text{if } \Delta v_i < 0 \\ 0 & \text{else} \end{cases}$$

and

$$\Delta v_i = \left(v_{i+\frac{1}{2}}^{n-\frac{1}{2}} - v_{i-\frac{1}{2}}^{n-\frac{1}{2}} \right)$$

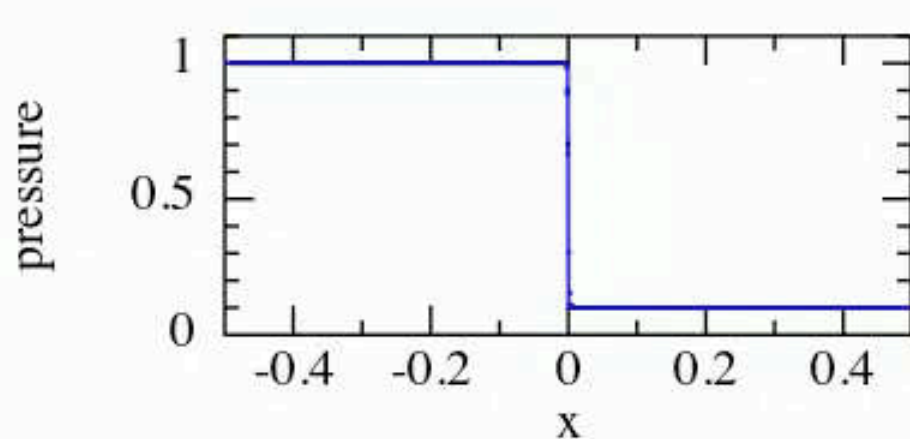
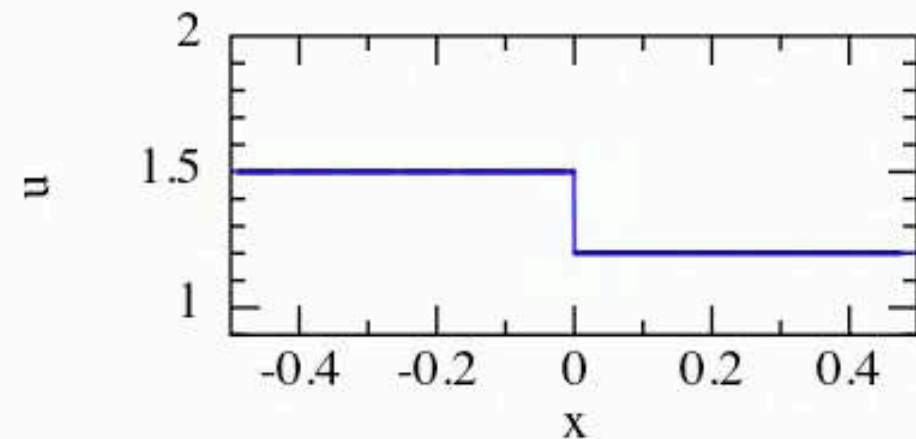
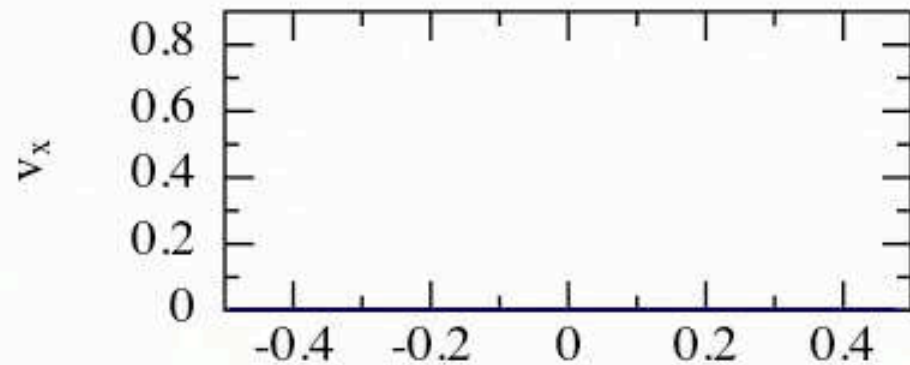
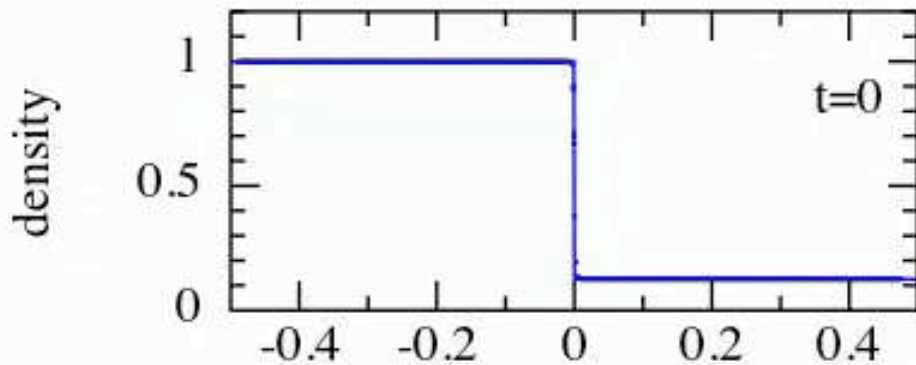
- Only add q for converging flows
- α is a value between 0 & 1, which can either be a fixed value, or dynamically calculated

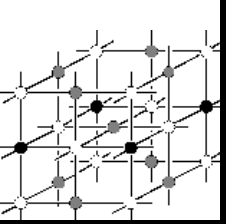


Sod Shock:

Artificial terms: Artificial viscosity

- Numerical algorithms are required for stability
- The form and parameterisation of these requires careful consideration to suppress numerical instabilities but not physical instabilities

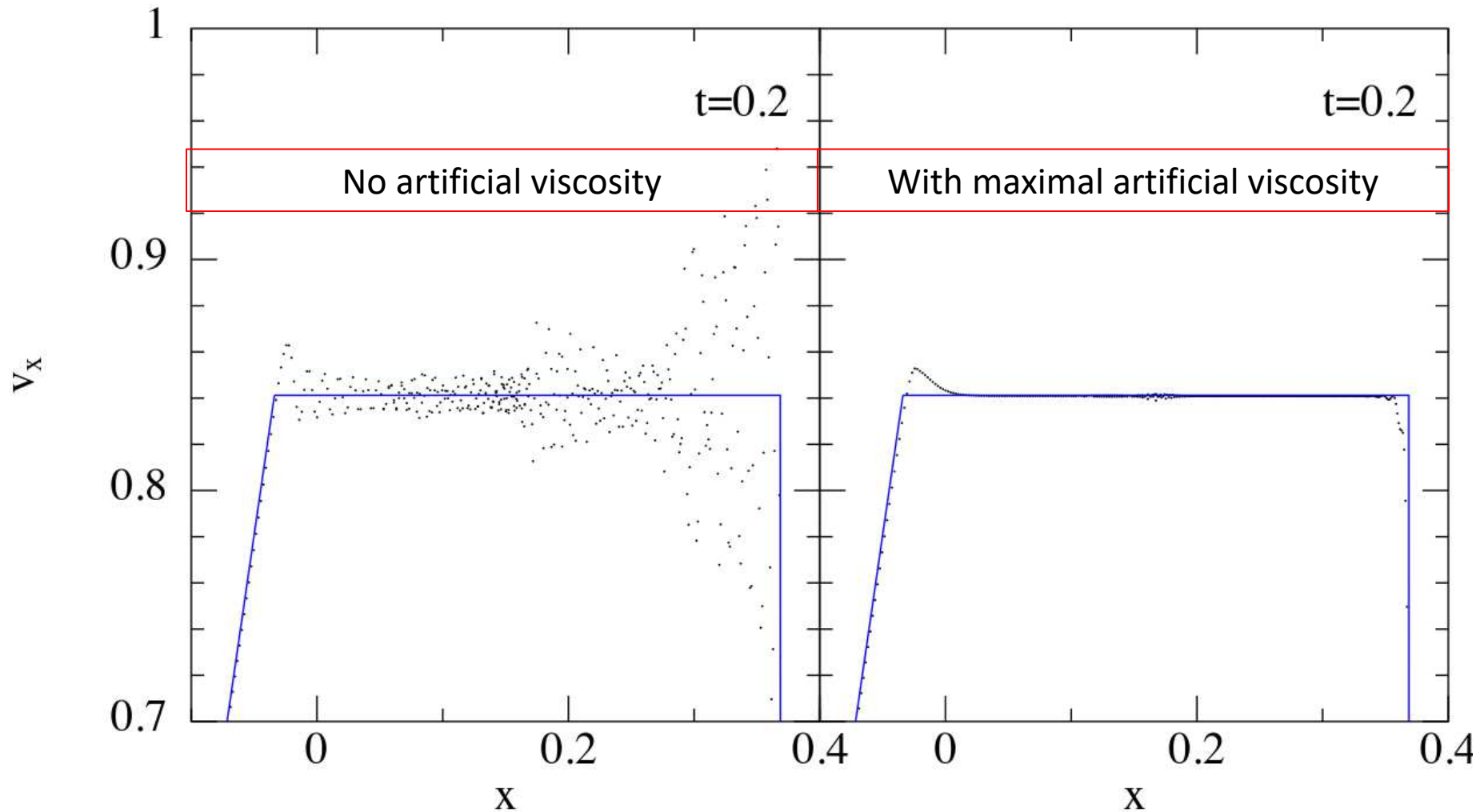


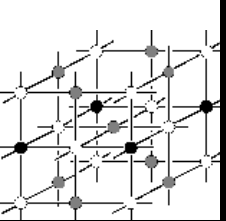


Sod Shock:

Artificial terms: Artificial viscosity

- Artificial viscosity well suppresses the ringing, and the numerical results (dots) better match the analytical result (blue)

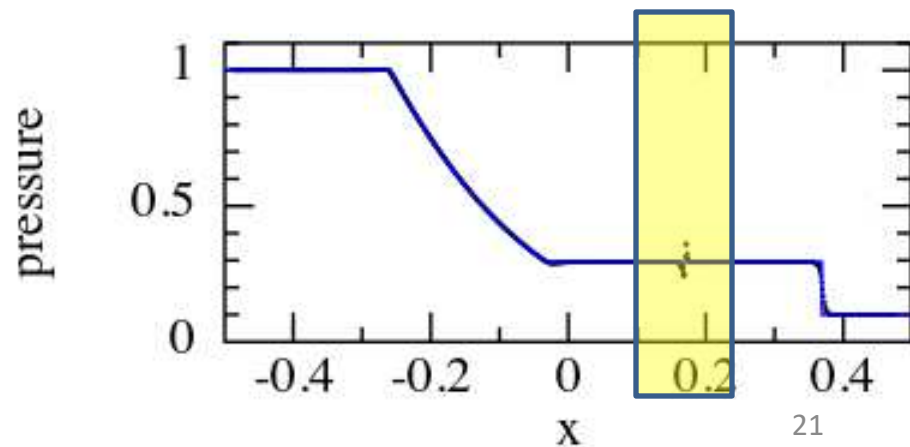
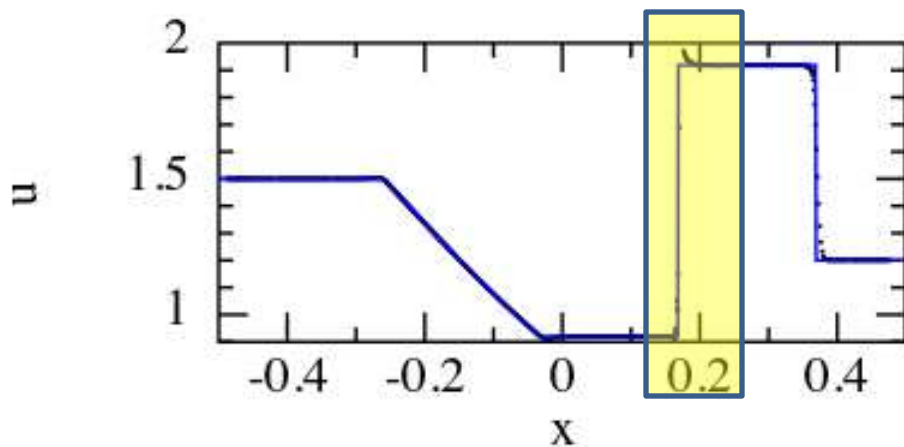
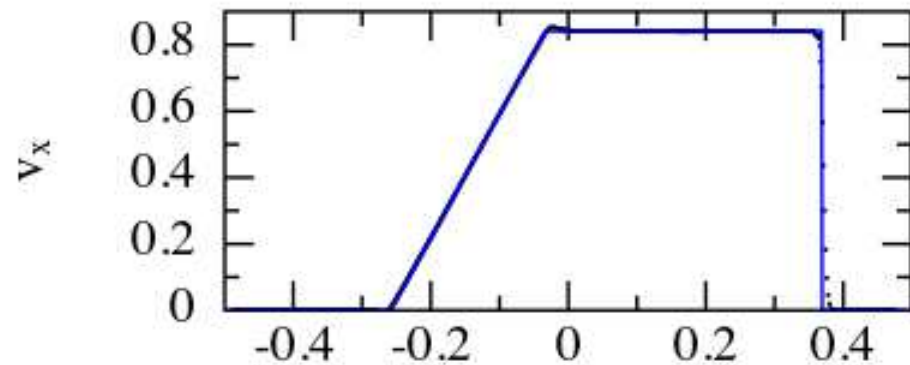
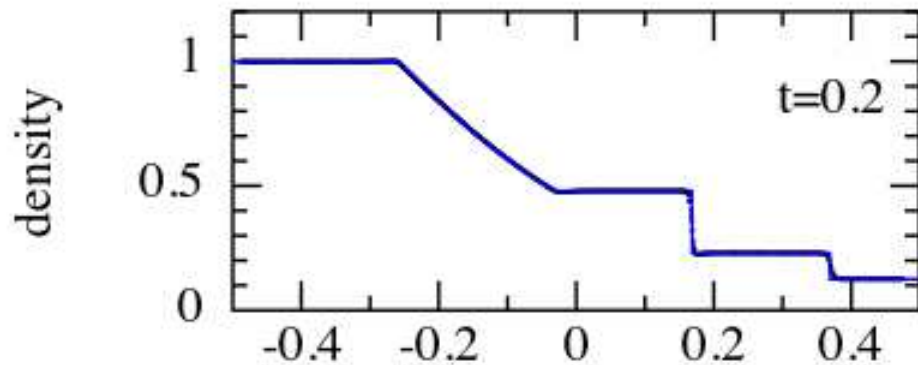


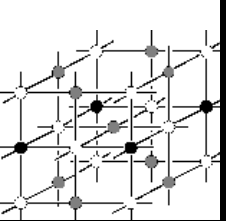


Sod Shock:

Artificial terms: Artificial viscosity

- Even with artificial viscosity, there is still some ‘blips’ in energy and pressure occurring at the **contact discontinuity**





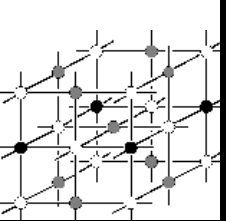
Sod Shock:

Artificial terms: Artificial conductivity

- Numerical algorithms are required for stability
- The form and parameterisation of these requires careful consideration to suppress numerical instabilities but not physical instabilities
- Since there is no instabilities in velocity at the contact discontinuity, we need to add an artificial term to the internal energy: artificial conductivity:

$$\frac{du}{dt} = -\frac{P}{\rho} \nabla \cdot \mathbf{v} + \left. \frac{du}{dt} \right|_{\text{artificial}}$$

- There are various forms of artificial conductivity throughout the literature



Sod Shock:

Artificial terms: Art. viscosity + art. conductivity

- Numerical algorithms are required for stability
- Alternatively, we can use

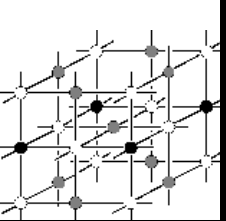
$$\frac{du}{dt} = -\frac{P + q}{\rho} \nabla \cdot \mathbf{v}$$

- Given the similarities between the energy equation and equation of motion, we can instead use

$$P_{i+\frac{1}{2},j+\frac{1}{2}}^{n+1} = (\gamma - 1) \rho_{i+\frac{1}{2},j+\frac{1}{2}}^{n+1} u_{i+\frac{1}{2},j+\frac{1}{2}}^{n+1} + q_{i+\frac{1}{2}}^n$$

- with

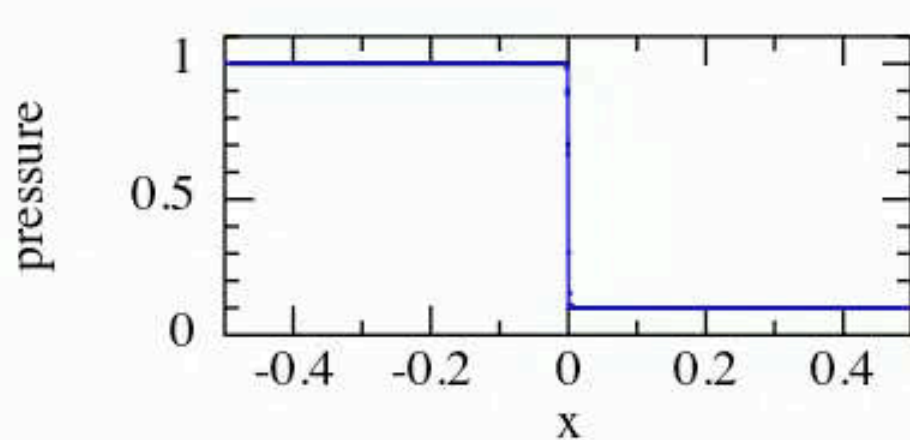
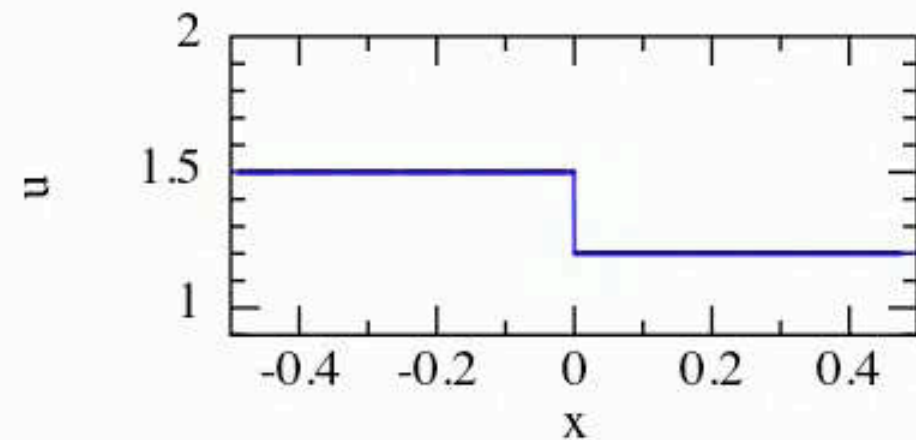
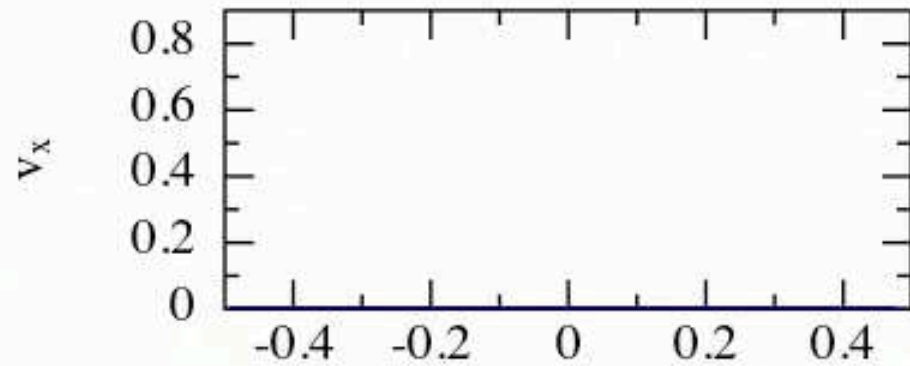
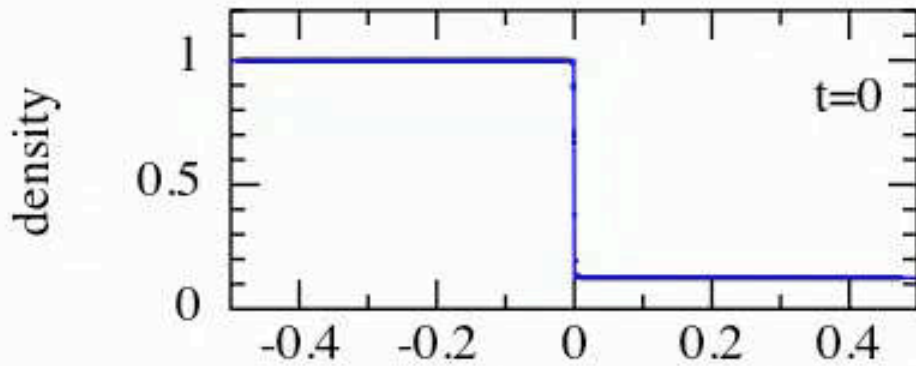
$$v_{x,i}^{n+\frac{1}{2}} = v_{x,i}^{n-\frac{1}{2}} - dt \left[\frac{2}{\rho_{i+\frac{1}{2}}^n + \rho_{i-\frac{1}{2}}^n} \frac{P_{i+\frac{1}{2}}^n - P_{i-\frac{1}{2}}^n}{dx} + v_{x,i}^{n-\frac{1}{2}} f(v) \right]$$
$$u_{i+\frac{1}{2}}^{n+1} = u_{i+\frac{1}{2}}^n - dt \left(\frac{P_{i+\frac{1}{2}}^n}{\rho_{i+\frac{1}{2}}^n} \frac{v_{x,i+1}^{n+\frac{1}{2}} - v_{x,i}^{n+\frac{1}{2}}}{dx} + \frac{v_{x,i+1}^{n+\frac{1}{2}} + v_{x,i}^{n+\frac{1}{2}}}{2} f(u) \right)$$

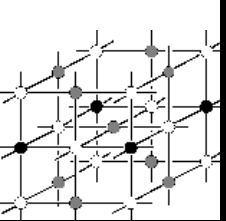


Sod Shock:

Artificial terms: Art. viscosity + art. conductivity

- Numerical algorithms are required for stability
- The form and parameterisation of these requires careful consideration to suppress numerical instabilities but not physical instabilities

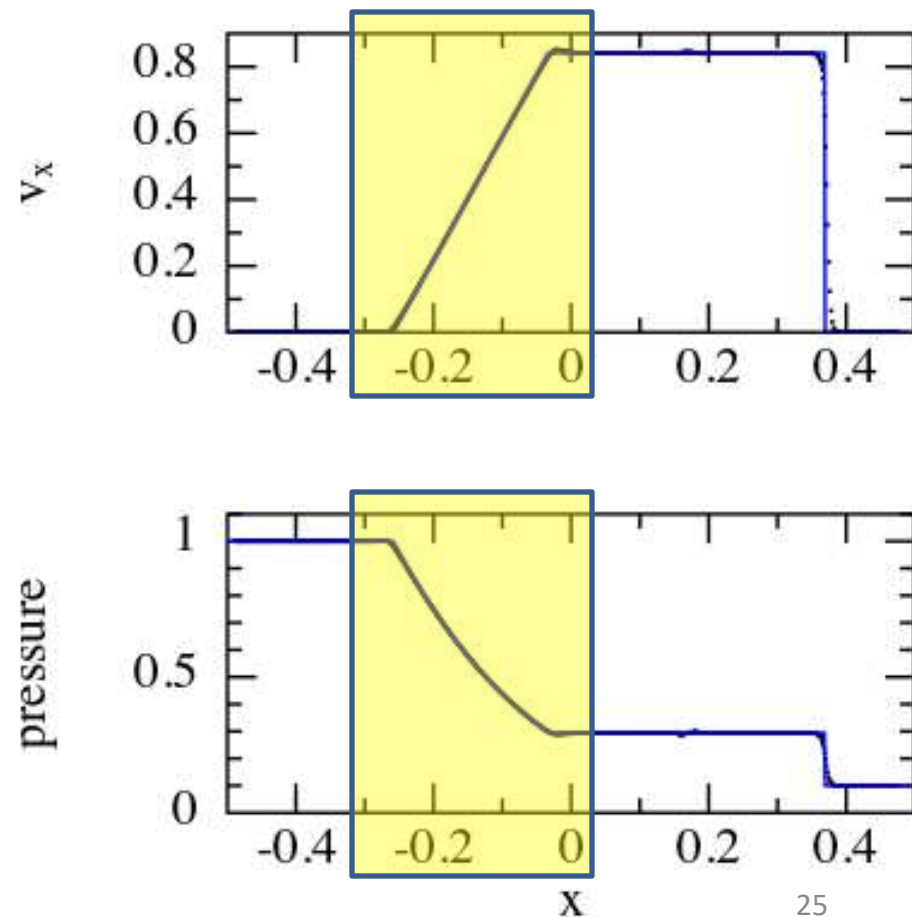
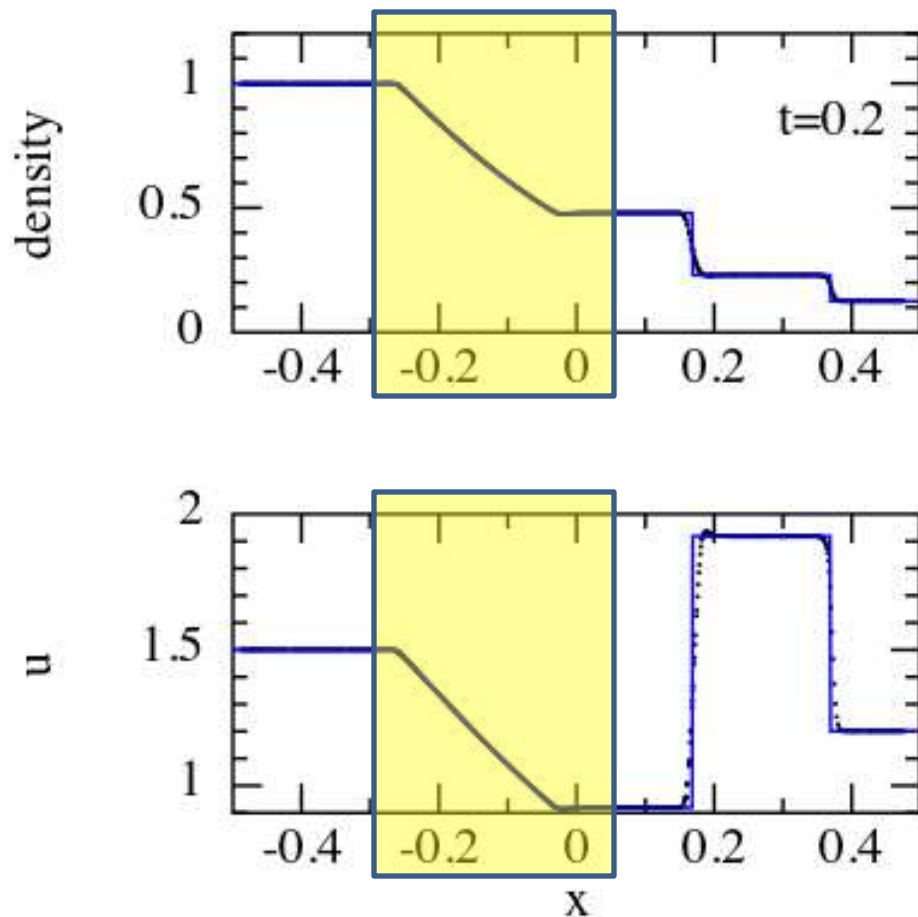


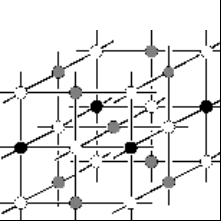


Sod Shock:

Artificial terms: Art. viscosity + art. conductivity

- The rarefaction wave is stable, thus we do not need to add in another numerical algorithms

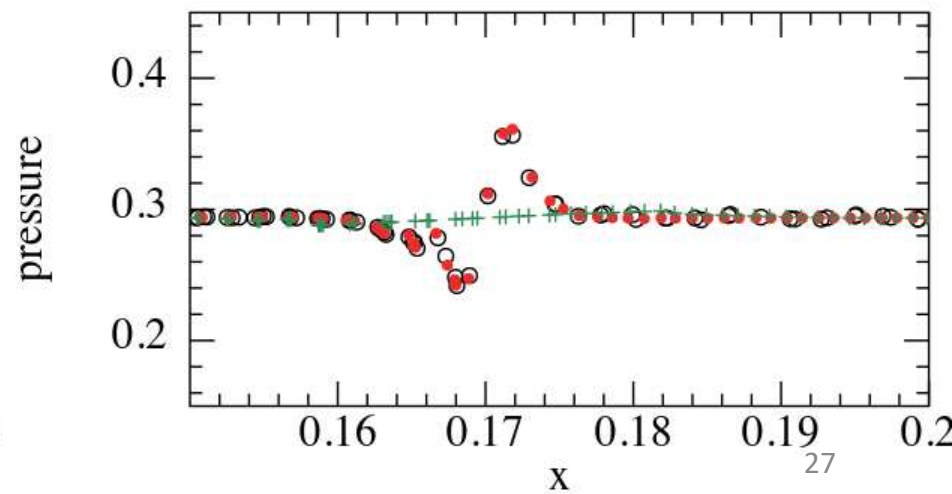
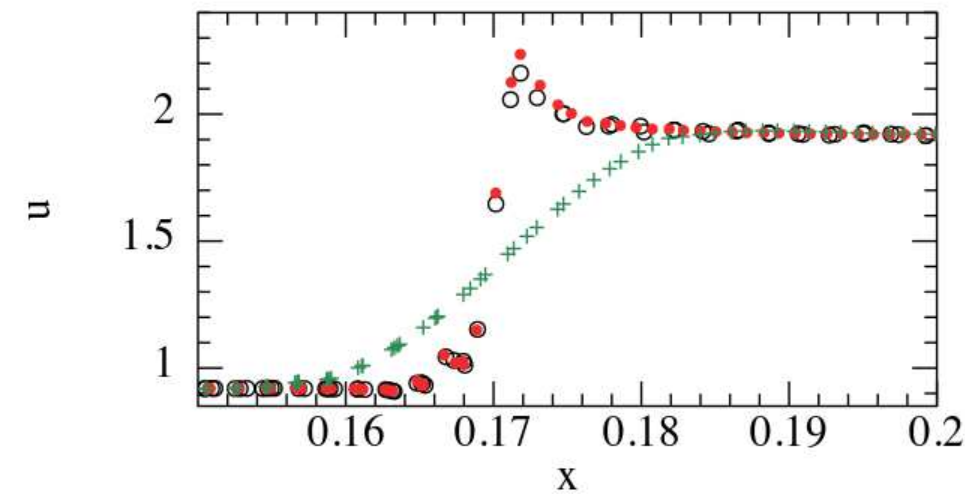
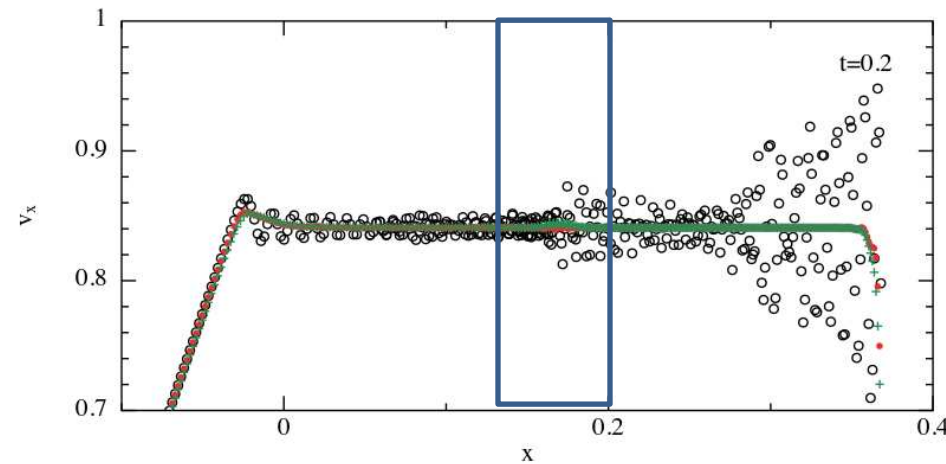
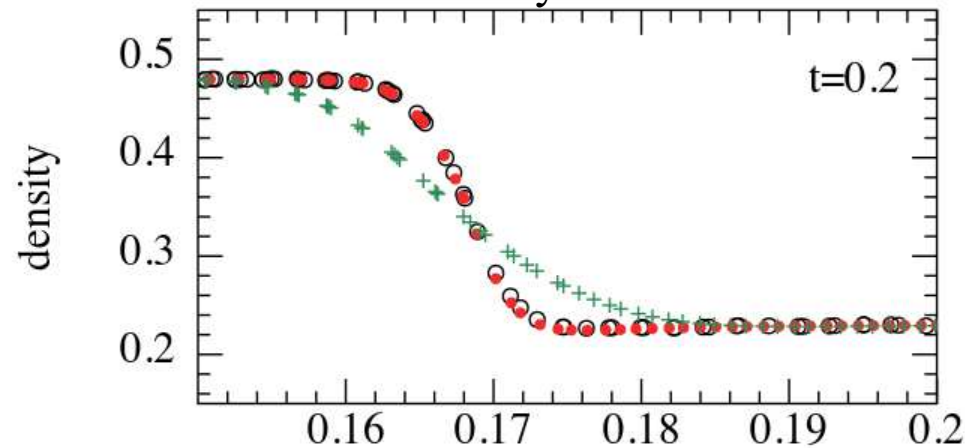


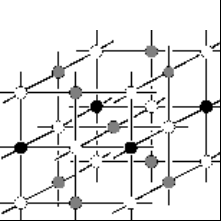


Sod Shock:

Artificial terms: Art. viscosity + art. conductivity

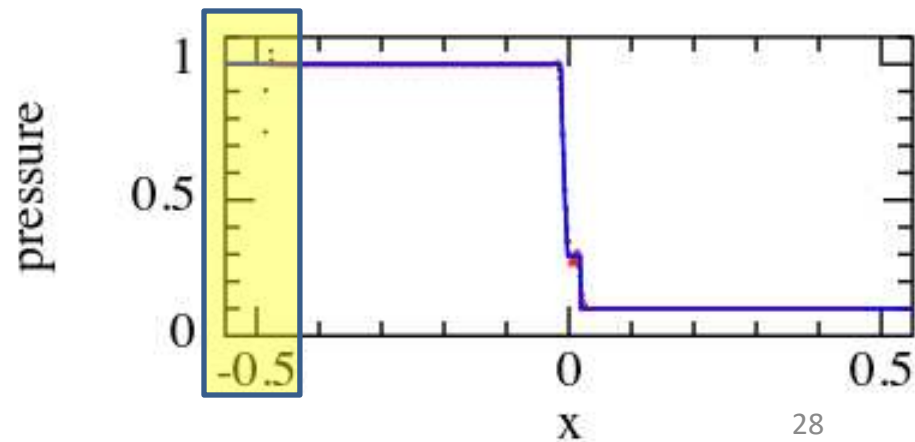
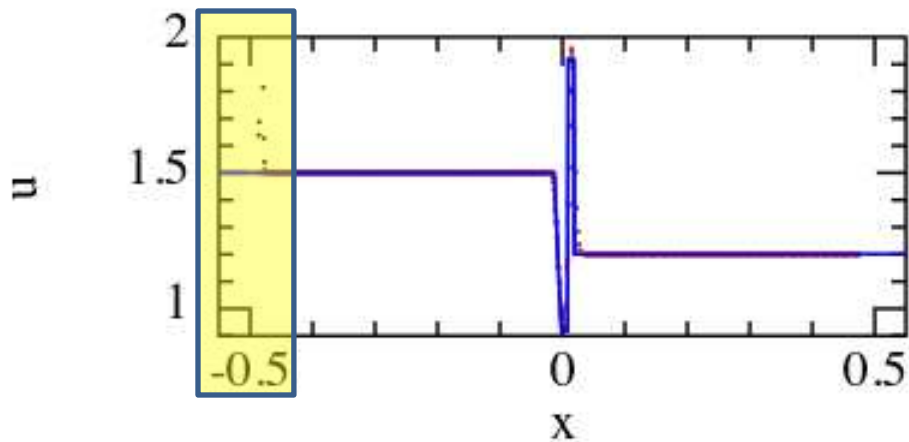
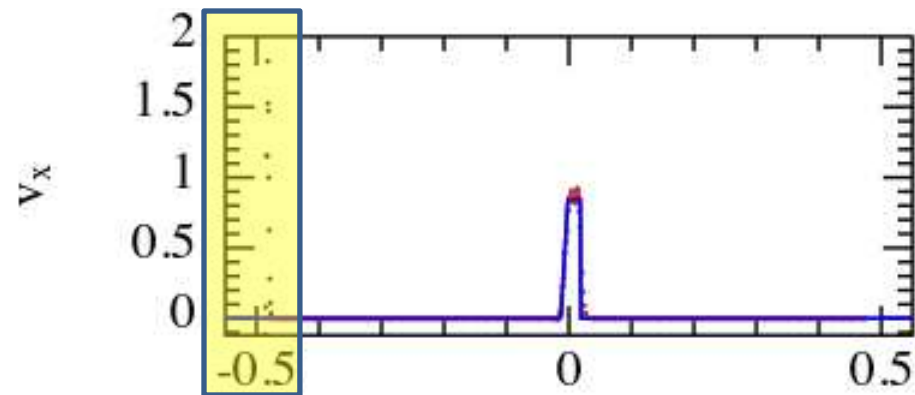
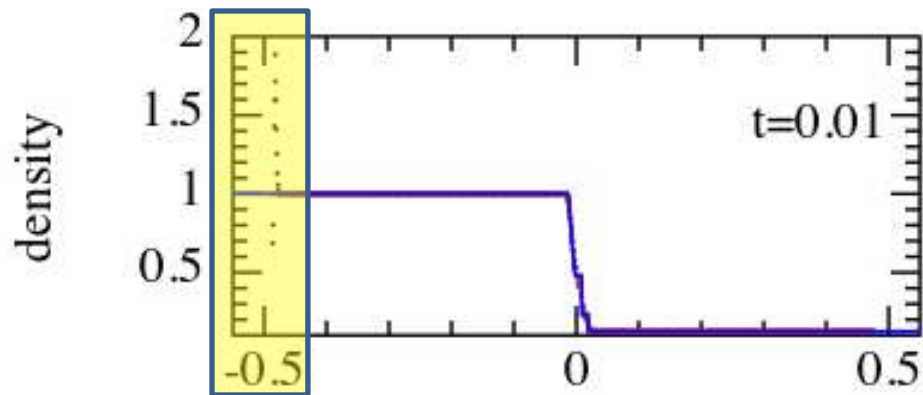
- Comparing without artificial terms (black), with artificial viscosity (red) & both (green) at the contact discontinuity

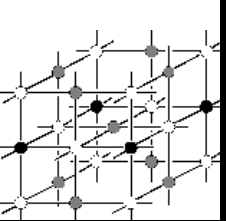




Sod Shock: Boundaries: A Cautionary Tale

- Incorrect boundaries (either by choice or a bug) will lead to incorrect answers

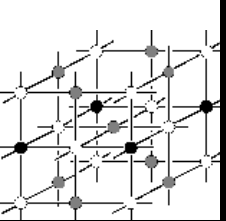




Conservation Laws

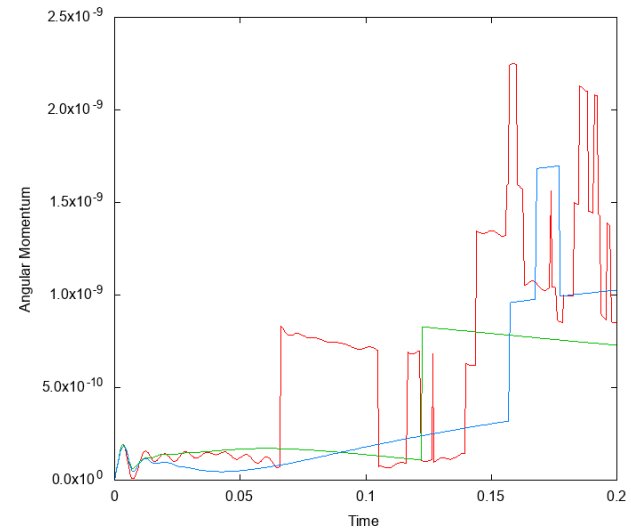
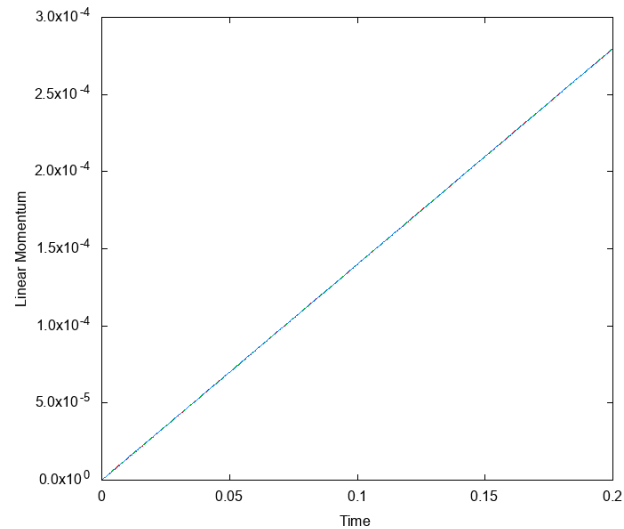
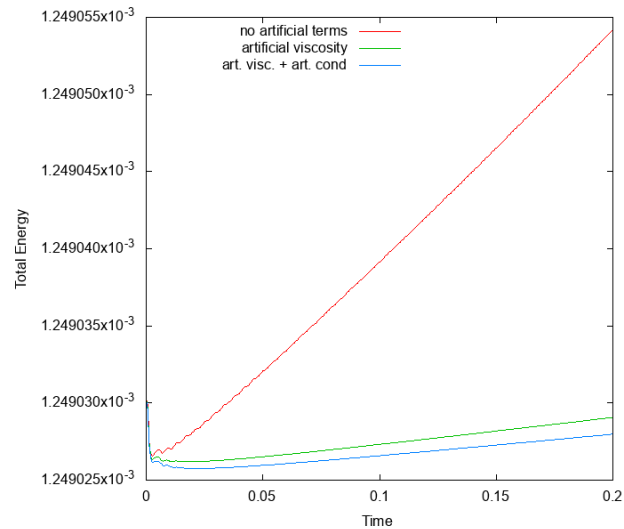
- Physics demands that certain quantities are conserved:
 - Mass
 - Energy
 - Linear momentum
 - Angular momentum

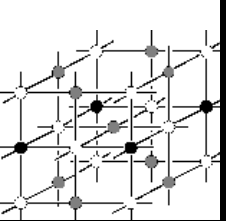
- Numerical experiments should also conserve these values
- Eulerian formalism (the equations in these slides) is not guaranteed to conserve mass
- SPH formalism (the graphs in these slides) is guaranteed to conserve mass by design



Conservation Laws

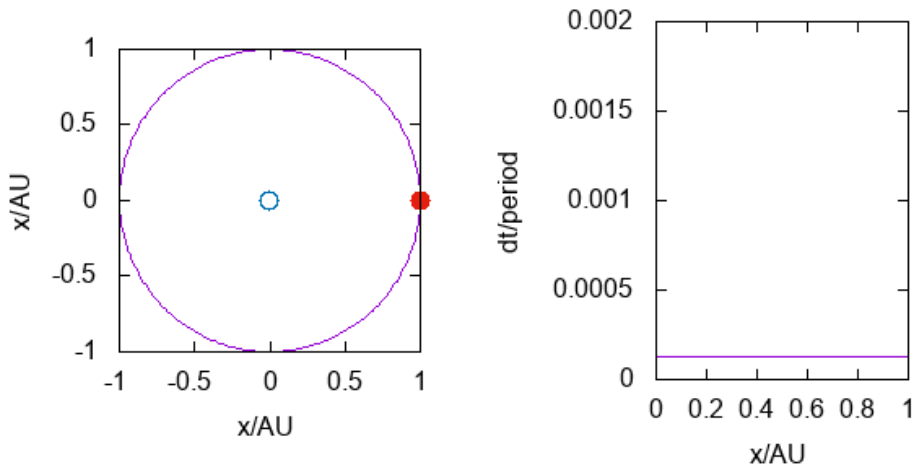
- Physics demands that certain quantities are conserved:
 - Mass
 - Energy
 - Linear momentum
 - Angular momentum
- Conserved quantities for the Sod Shock tube:



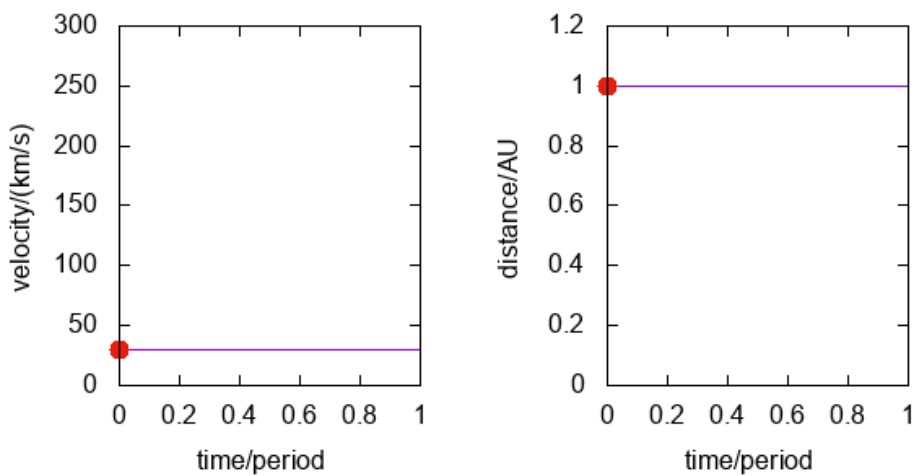


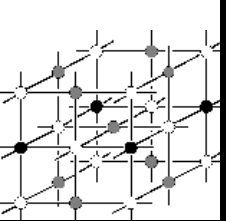
Timestepping

- As the simulation evolves, what timestep do we choose?
 - As long as possible, but short enough to resolve the physics
 - We want this to be chosen by the programme and not as an input parameter



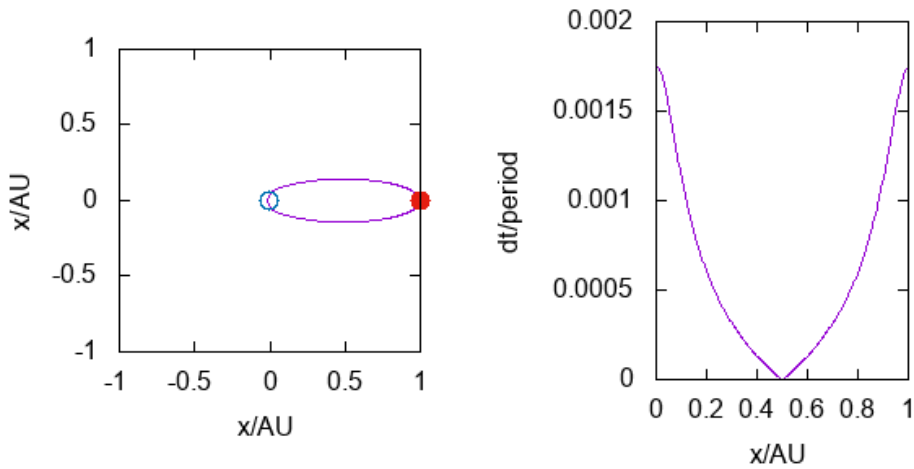
$$dt = \frac{1}{N} \frac{2\pi \vec{r}}{\vec{v}} \quad \text{where } N \text{ is the number of steps per orbit}$$
$$= C \frac{\vec{r}}{\vec{v}} \quad \text{where } C \leq 1$$



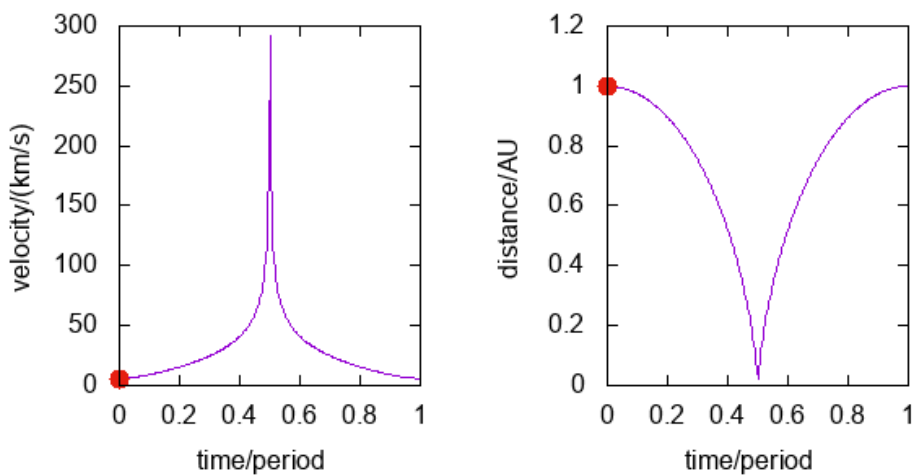


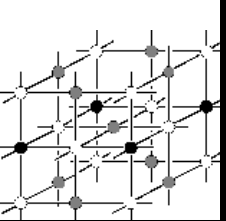
Timestepping

- As the simulation evolves, what timestep do we choose?
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 - We want this to be chosen by the programme and not as an input parameter



$$dt = \frac{1}{N} \frac{2\pi \vec{r}}{v} \quad \text{where } N \text{ is the number of steps per orbit}$$
$$= C \frac{\vec{r}}{v} \quad \text{where } C \leq 1$$





Timestepping

- The previous timestep can be extended to numerical fluid dynamics
- Courant–Friedrichs–Lewy is a common limiting timestep:

$$dt = \min \left(C \frac{dx_i}{v_i}, C \frac{dx_i}{c_{s,i}} \right) \quad \text{where} \quad C \leq 1$$

$\longleftarrow dx_i \longrightarrow$

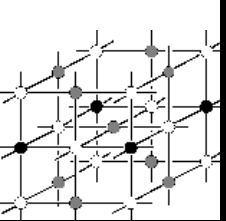


$v_{x,i-1}$

$v_{x,i}$

$v_{x,i+1}$

- This is based upon “How long does it take information (e.g. a wave) to travel from one side of the cell to the other?”
- If a wave travels too fast, then it is unresolved and the simulation may break



Timestepping

A practical warning

- Although testable on well known problems, selecting a dt coefficient can be challenging in practical simulations
- A timestep governing non-ideal MHD is

$$dt = C \frac{dx_i^2}{\eta_i}$$

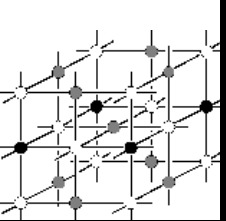
where

$$C = 1/6 \text{ (Bai 2014)}$$

$$C = 1/2\pi \text{ (Wurster+ 2016)}$$

$$C = 1/4\pi \text{ (Wurster, miscellaneous tests)}$$

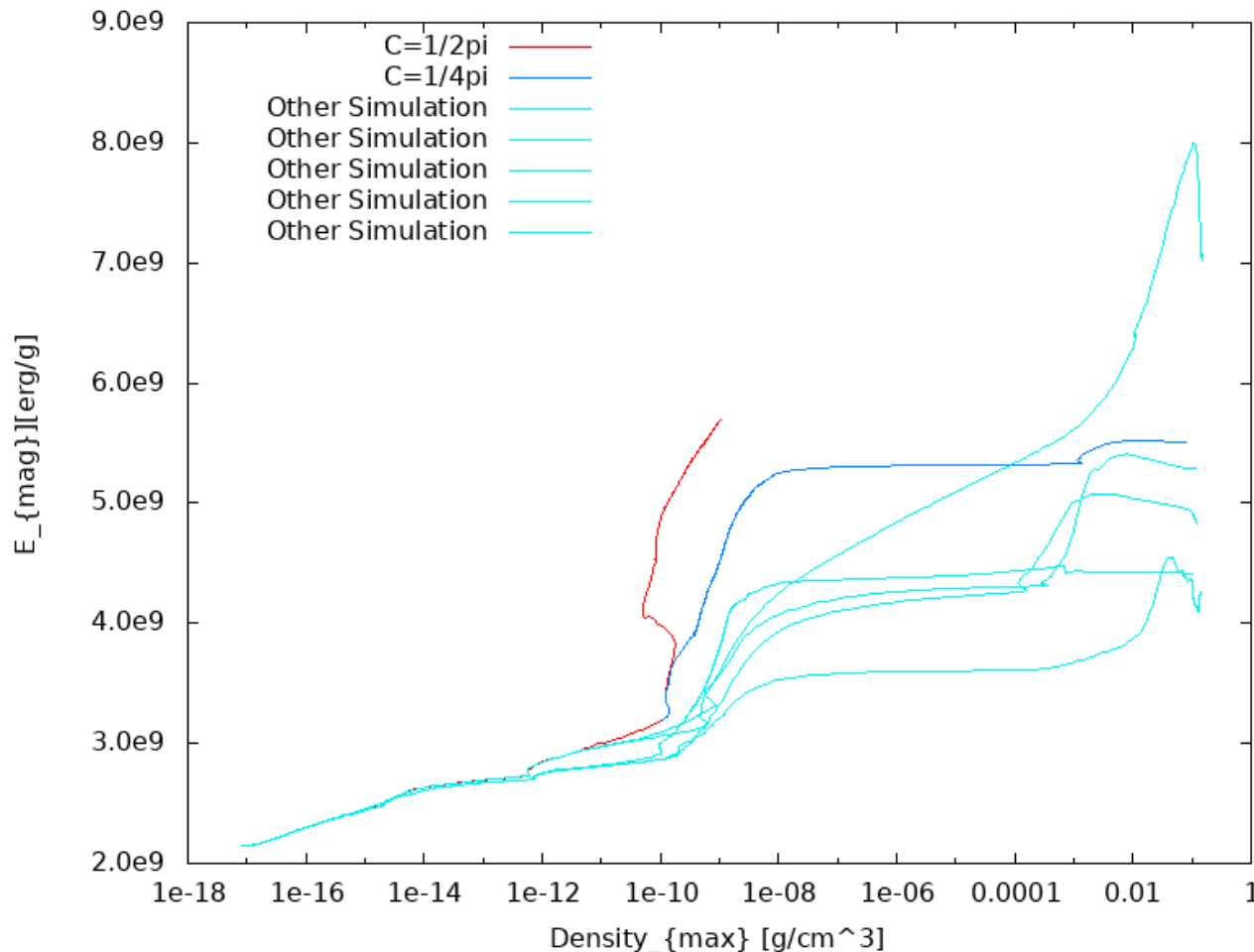
$$C = 1/10\pi \text{ (Tsukamoto+2015)}$$

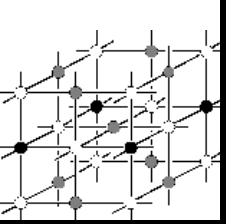


Timestepping

A practical warning

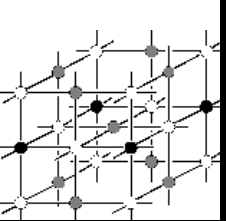
- Although testable on well known problems, selecting a dt coefficient can be challenging
- Each line represent a different simulation. The red line does not match the rest of the trends
- The blue line is the same model, but with a smaller C



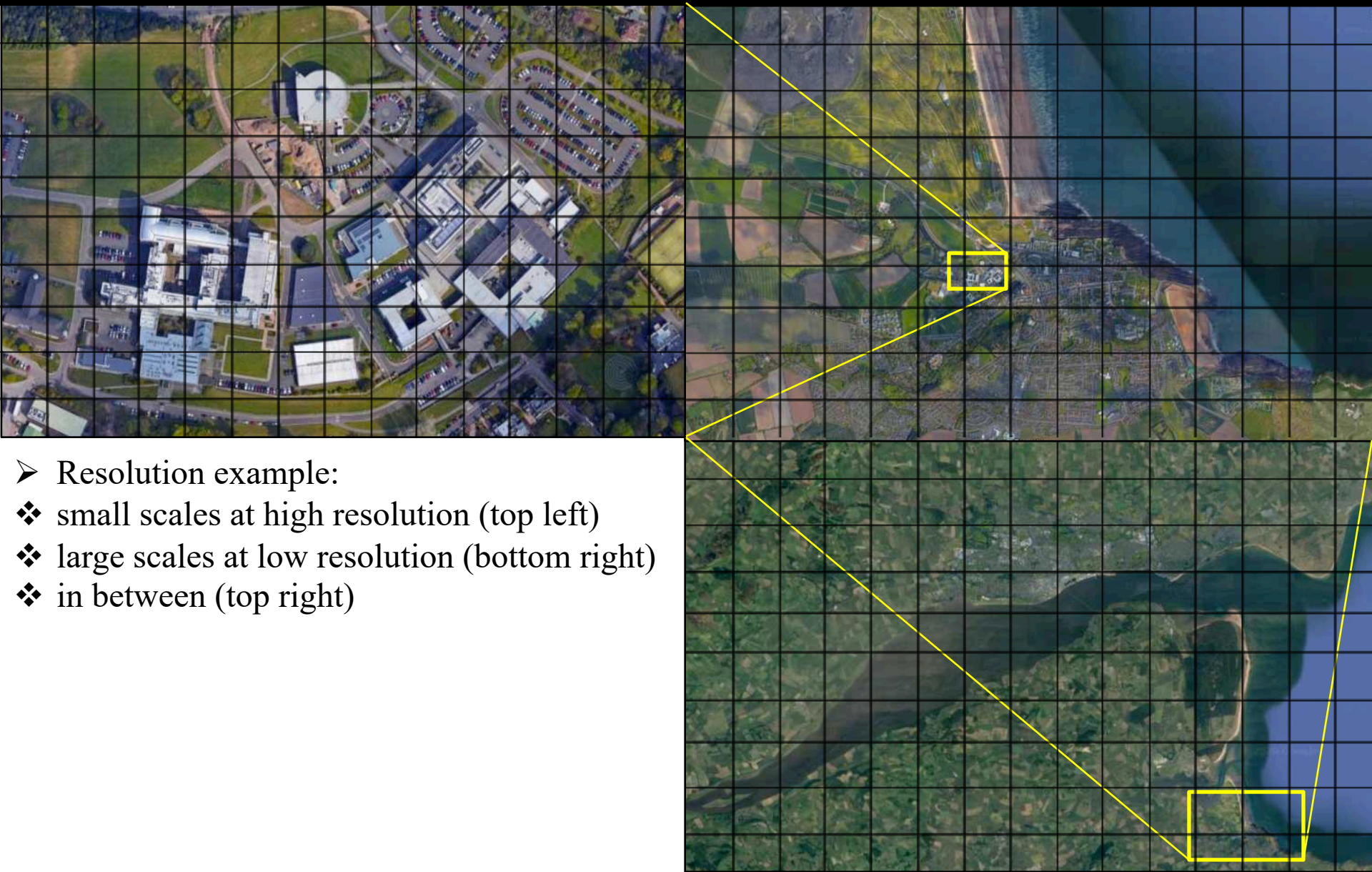


Resolution

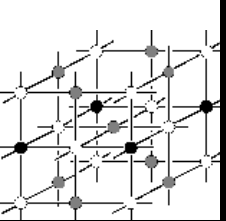
- Computers have finite resources, and users have finite time/patience
- To determine the desired resolution, must carefully consider
 - What are you trying to model?
 - What are your computational resources?
- Our options are generally
 - Model a small region at high resolution
 - Model a large region at low resolution



Resolution: Geographic example

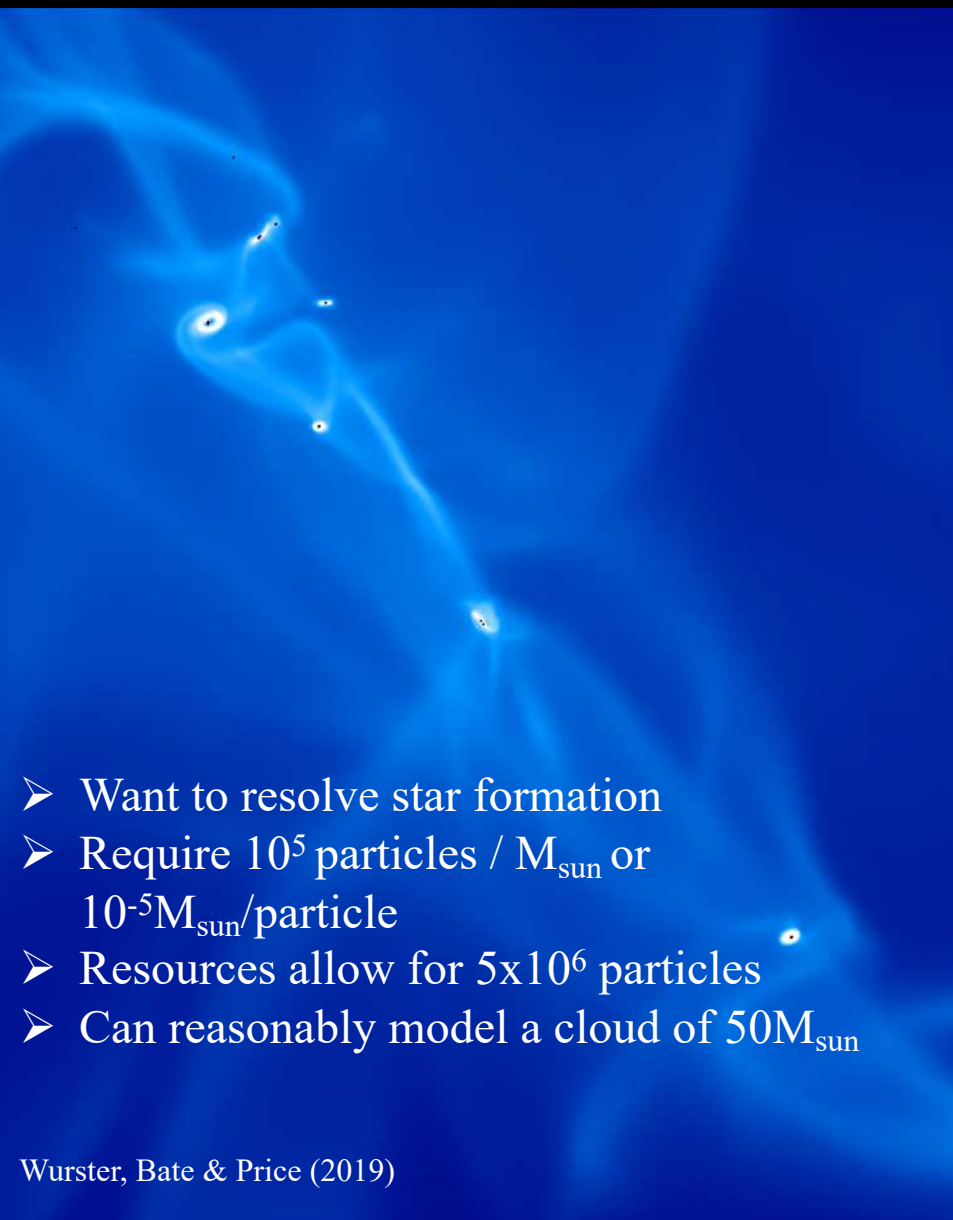


- Resolution example:
- ❖ small scales at high resolution (top left)
- ❖ large scales at low resolution (bottom right)
- ❖ in between (top right)



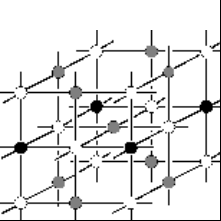
Resolution:

Example: Modelling a molecular cloud



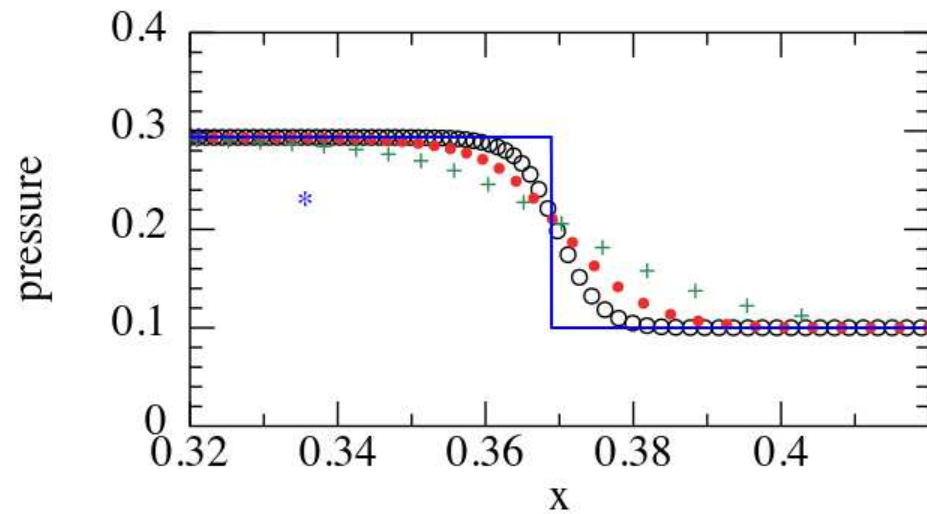
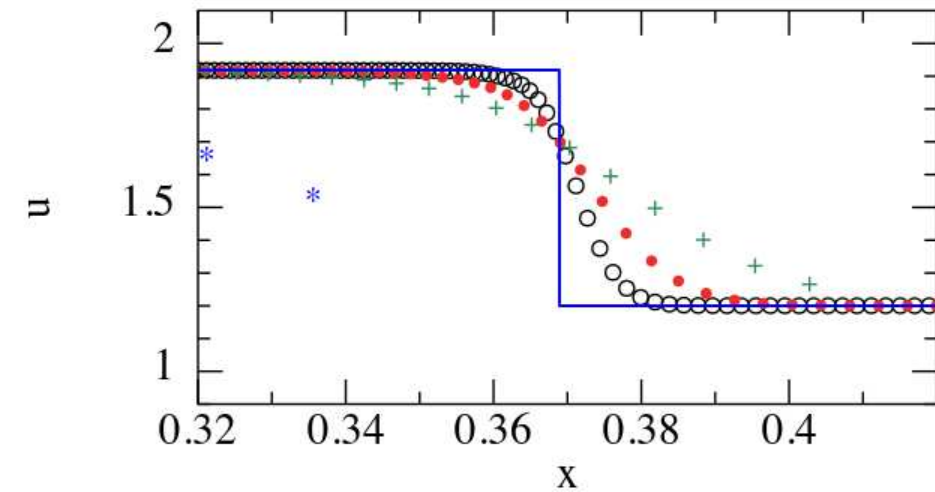
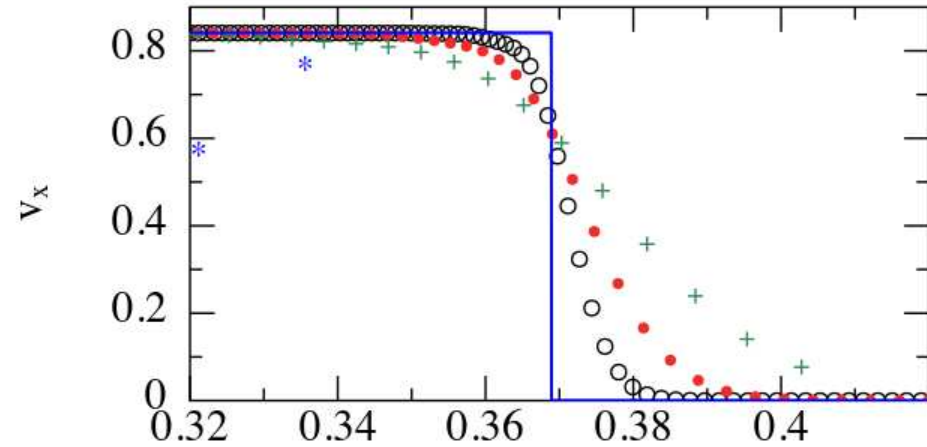
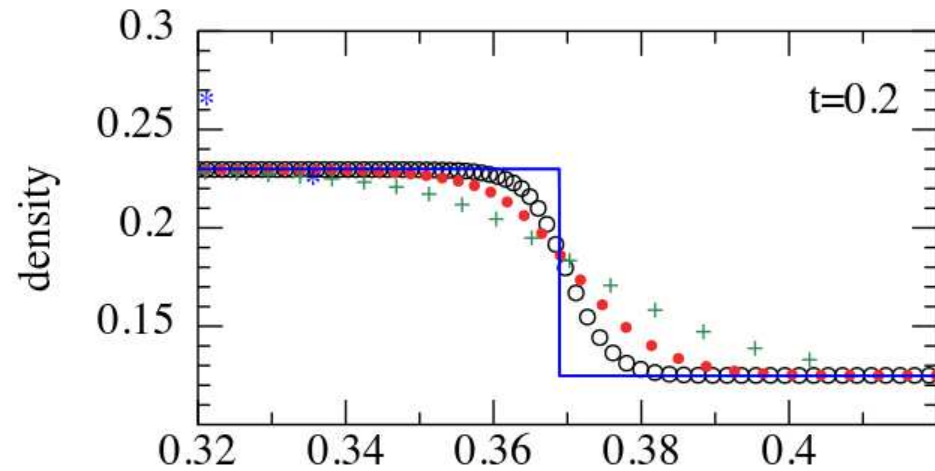
- Want to resolve star formation
- Require 10^5 particles / M_{sun} or $10^{-5}M_{\text{sun}}/\text{particle}$
- Resources allow for 5×10^6 particles
- Can reasonably model a cloud of $50M_{\text{sun}}$

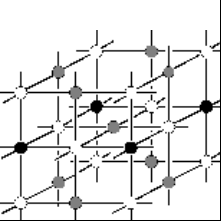
- Want to model an entire molecular cloud
- Cloud contains $10^5 M_{\text{sun}}$
- Resources allow for 5×10^6 particles
- Can reasonably resolve down to $0.02M_{\text{sun}}$



Resolution: Sod Shock

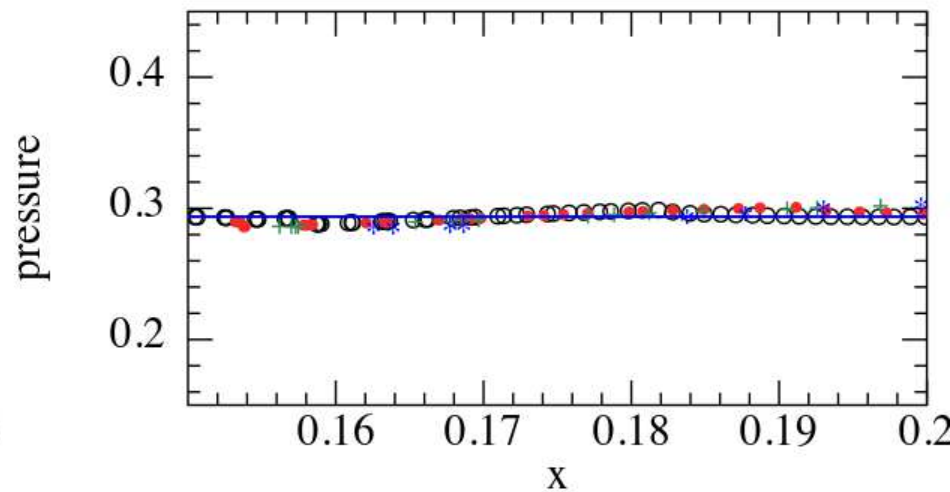
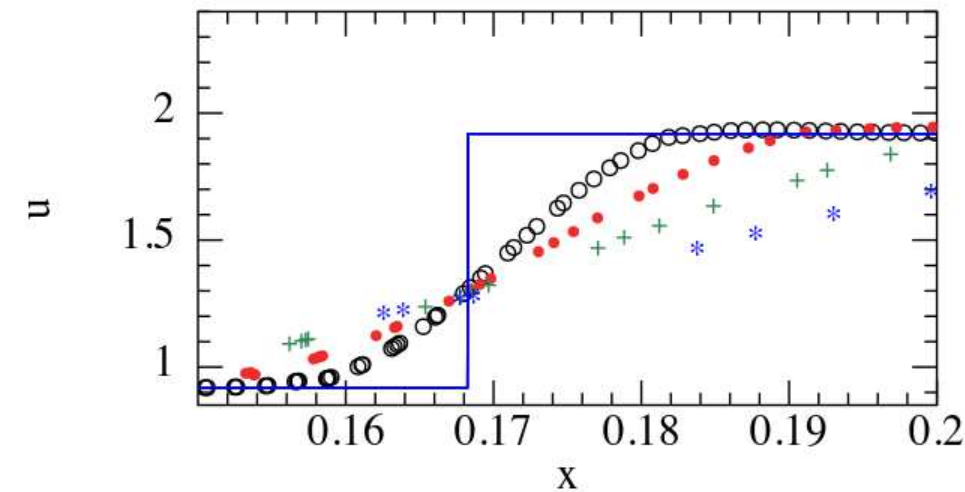
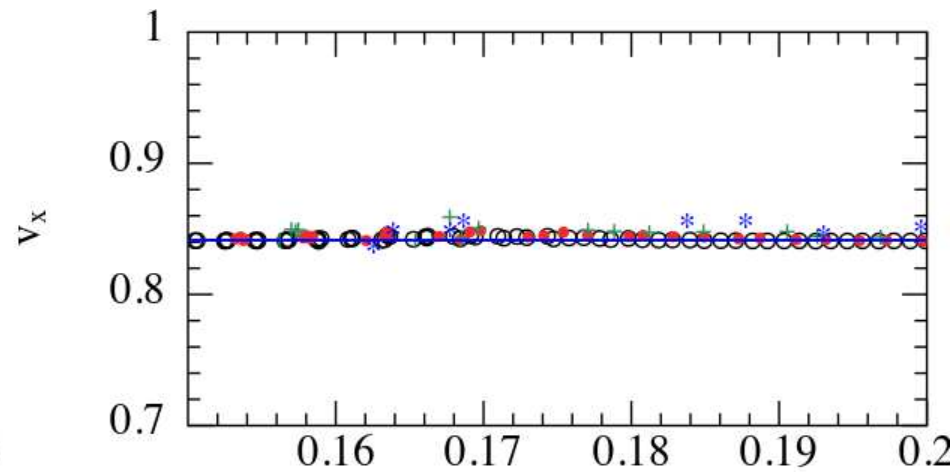
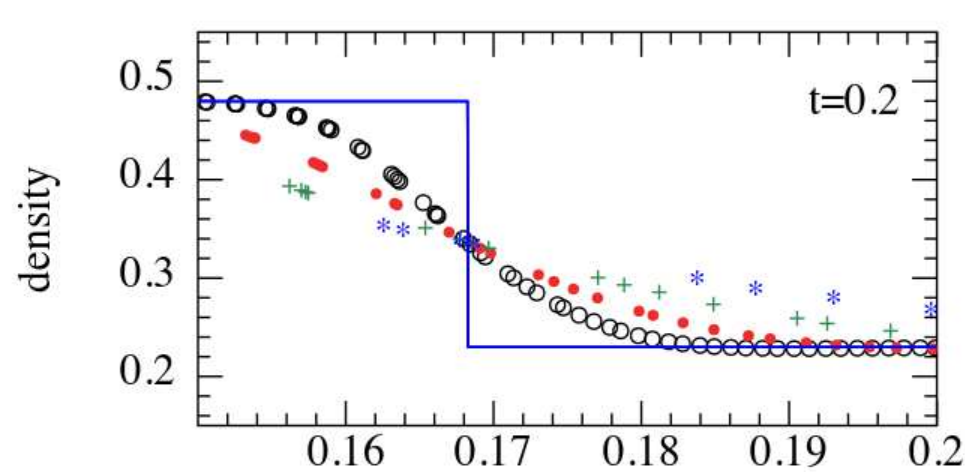
- Investigating the shock wave for four resolutions: $n_{x,\text{left}} = 32, 64, 128$ & 256

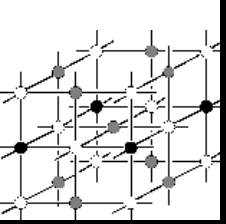




Resolution: Sod Shock

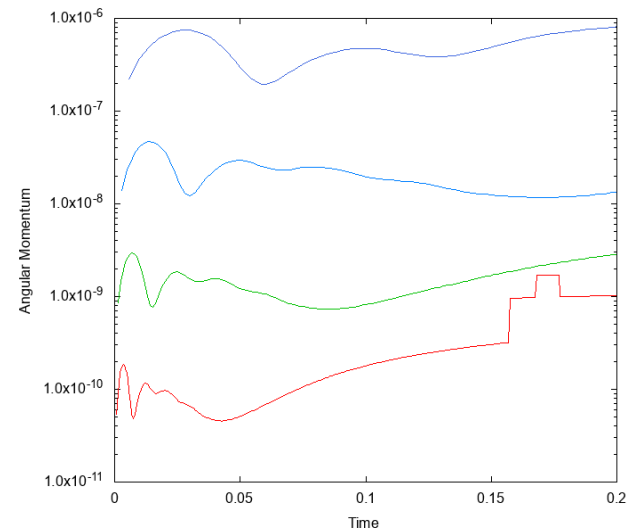
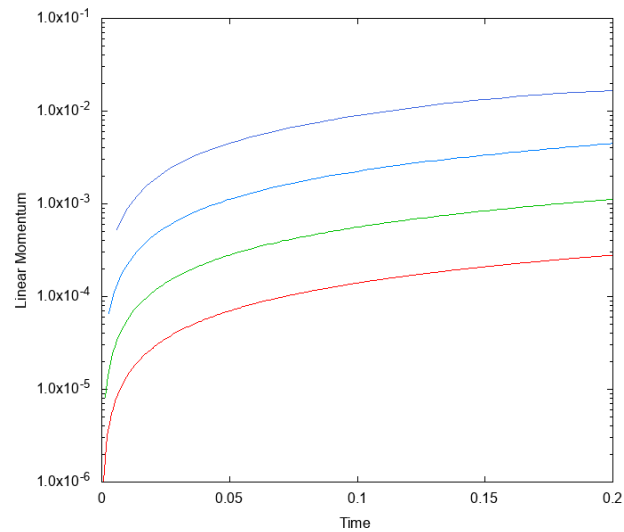
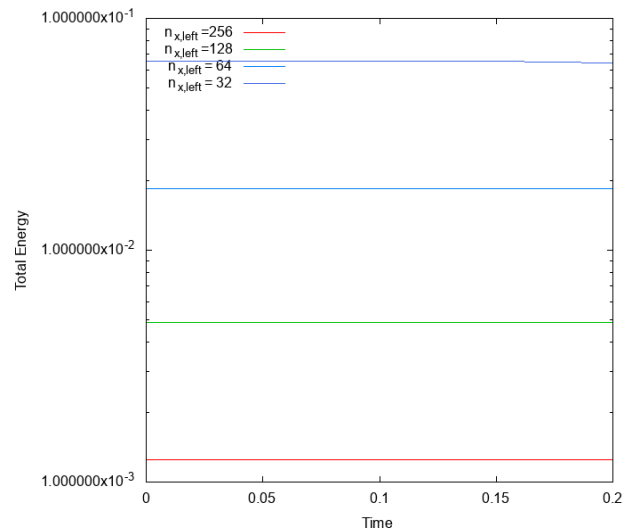
- Investigating the contact discontinuity for four resolutions: $n_{x,\text{left}} = 32, 64, 128$ & 256

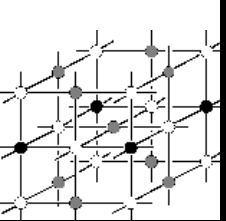




Resolution: Sod Shock

- Testing four resolutions: $n_{x,\text{left}} = 32, 64, 128$ & 256
- Conserved quantities get better with increasing resolution





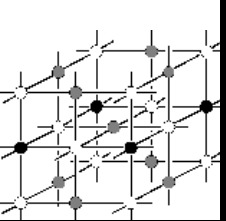
Resolution: Warning!

- Recall: decreasing dx by a factor 2
 - doubles the number of calculations per step
 - doubles the number of steps

$$dt = C \frac{dx_i}{v_i} \quad \text{where} \quad C \leq 1$$

In numerical studies, the user must always balance resolution with runtime!





Resolution: Sod Shock

- Testing four resolutions: $n_{x,\text{left}} = 32, 64, 128$ & 256
- Runtime is considerably longer for higher resolutions

